**Geoscientific
Model Development**

The chemical transport model Oslo CTM3

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Abstract. We present here the global chemical transport model Oslo CTM3, an update of the Oslo CTM2. The update comprises a faster transport scheme, an improved wet scavenging scheme for large scale rain, updated photolysis rates and a new lightning parameterization. Oslo CTM3 is better parallelized and allows for stable, large time steps for advection, enabling more complex or high spatial resolution simulations. A new treatment of the horizontal distribution of lightning is presented and found to compare well with measurements. The vertical distribution of lightning is updated and found to be a large contributor to CTM2–CTM3 differences, producing more NO_x in the tropical middle troposphere, and less at the surface and at high altitudes. Compared with Oslo CTM2, Oslo CTM3 is faster, more capable and has better conceptual models for scavenging, vertical transport and fractional cloud cover. CTM3 captures stratospheric O_3 better than CTM2, but shows minor improvements in terms of matching atmospheric observations in the troposphere. Use of the same meteorology to drive the two models shows that some features related to transport are better resolved by the CTM3, such as polar cap transport, while features like transport close to the vortex edge are resolved better in the Oslo CTM2 due to its required shorter transport time step. The longer transport time steps in CTM3 result in larger errors, e.g., near the jets, and when necessary the errors can be reduced by using a shorter time step. Using a time step of 30 min, the new transport scheme captures both large-scale and small-scale variability in atmospheric circulation and transport, with no loss of computational efficiency. We present a version of the new transport scheme which has been specifically tailored for polar studies, resulting in more accurate polar cap transport than the standard CTM3 transport, confirmed by comparison to satellite observations. Inclusion of tropospheric sulfur chemistry and nitrate aerosols in CTM3 is shown to be important to reproduce tropospheric O_3, OH and the CH_4 lifetime well.

1 Introduction

The University of Oslo chemistry-transport model, Oslo CTM2, has been used extensively over the past decade for studies of stratospheric and tropospheric chemistry, greenhouse gases, and climate forcing (Wild et al., 2003; Gauss et al., 2003, 2006; Berglen et al., 2004; Isaksen et al., 2005; Dalsøren et al., 2007; Solberg et al., 2008; Søvde et al., 2008, 2011a; Hoor et al., 2009; Dalsøren et al., 2010; Myhre et al., 2011; Hodnebrog et al., 2011). CTM2 resulted from a collaboration between Oslo and the University of California Irvine (UCI) in which the UCI development of CTMs (Prather et al., 1987; Hall and Prather, 1993, 1995) was combined with the Oslo development of forecast meteorology fields from the European Centre for Medium-Range Weather Forecasts (ECMWF) (Sundet, 1997). The Oslo CTM2 was first documented in Jonson et al. (2001). Since CTM2, better and more efficient transport methods and diagnostics have been developed (Prather et al., 2008, 2011), and more accurate simulation of atmospheric processes, such as photolysis and scavenging, based on the ECMWF forecast data (Neu et al., 2007; Neu and Prather, 2012). These developments have been merged with CTM2's stratosphere-plus-troposphere, gas-plus-aerosol, chemistry model to form the Oslo CTM3, which is documented and evaluated here.

Section 2 describes the new components of Oslo CTM3, noting the main differences from CTM2. We focus on transport and the gas-phase chemistry components, so our

simulations will also include aerosol modules that affect chemistry directly. Section 3 evaluates the new model against some traditional chemistry climatologies and new observational case studies. We show that the inclusion of the tropospheric sulfur cycle (Berglen et al., 2004) and nitrate aerosols (Myhre et al., 2006), which directly affect the chemistry, is important for tropospheric chemistry. Also wet scavenging and transport sensitivity studies are evaluated in Sect. 3, while Sect. 4 concludes the study.

2   Model description

The Oslo CTM3 is a global 3-D CTM, driven by 3-hourly meteorological forecast data from the European Centre for Medium-Range Weather Forecasts (ECMWF) Integrated Forecast System (IFS) model, produced on a daily basis with 12 h of spin-up starting from an analysis at noon on the previous day, which are then pieced together to give a uniform dataset for a whole year. It has been clear in the history of ECMWF forecasts that there is a spin-up time for the forecasts during which short-lived phenomena like precipitation adjust to inadequate initialization from the analysis fields (Kraabøl et al., 2002; Wild et al., 2003). Forecasts also allow retrieval of additional meteorological data such as convective mass fluxes, which are not available from operational analyses. The 12-hour spin-up used here was based on ECMWF experience. The use of pieced forecasts is also imperative when running an externally forced CTM because much of the physics need to be integrated, not snapshots (e.g., 3-D clouds, convective fluxes, and precipitation fields), which are integrated from the forecast model. If analysis fields are used as the core meteorology in the CTM, then the CTM must include the full general circulation model physics (unlike here). Winds, temperature, pressure and humidity are given as instantaneous values every 3 h, while other variables such as rainfall and convective fluxes are averages over the 3-h interval. The cycle used here is 36r1, starting from ERA-Interim re-analyses.

The 60-layer vertical resolution of the IFS is used in CTM3. The resolution near the tropical tropopause is ~ 1 km, and the uppermost, 10-km thick layer is centered at 0.11 hPa (about 60 km). A horizontal Gaussian-grid of resolution T42 ($\sim 2.8° \times \sim 2.8°$) is the standard resolution of CTM3, but the original resolution of the IFS (T319L60, $\sim 0.5° \times \sim 0.5°$) can also be used. The UCI/Oslo CTM framework is flexible and can be used on any 3-D quadrilateral grid.

All modelled processes, or operations, are carried out sequentially and asynchronously with the only requirement that all process sub-cycles must synch at the end of the operator-split time step, which is typically 60 min but can be shorter or longer. As will be explained in Sect. 2.2.1, the large-scale transport selects a maximum global time step from the Lifshitz criterion, and then sets the number of sub-

Table 1. The main model updates documented in this work.

a	A new advection core that greatly speeds up the model, enabling more complex chemistry or very-high resolution (T319) chemistry simulations (Prather et al., 2008).
b	New, modernized lightning NO_x parameterizations that reflect current satellite observations and recent campaigns (documented here).
c	New scavenging scheme that includes liquid and ice water and partially overlapping clouds (Neu and Prather, 2012).
d	New fast-JX combined with the cloud overlap scheme (Neu et al., 2007) with an additional speed up from Neu's quadrature to a randomly selected cloud profile selected hourly based on the cloud fractions and overlaps from the ECMWF forecast data.
e	New convection treatment using ECMWF-diagnosed entrainment and detrainment, along with combining the convection and large-scale vertical advection to save time and reduce noise in redundant vertical transport where the inferred convective subsidence nearly cancels the large-scale convergence (Prather et al., 2008).

cycle steps, e.g., within 60 min, meeting that criterion. For emissions, boundary layer mixing, chemistry and dry deposition (called the EBCD-sequence below), CTM3 retains the CTM2 method, with an internal cycling of maximum 15 min. If a 60 min sub-step is set for the EBCD-sequence in an operator-split time step of 60 min, the EBCD-sequence is carried out 4 times with 15 min time step. Note that the time step used in chemical integrations may be shorter than this (see Sect. 2.1), and does not change unless the operator-split time step is shorter than 15 min. As an example, suppose the Lifshitz maximum time step is 42 min, then two 30 min sub-cycle advection steps will be calculated. If e.g. the EBCD-sequence is calculated every 10-min step and transport and wet scavenging request a 15-min step, then the sub-cycle picks 1/12 of 60 min, or 5 min, as its basic cycle. At 5 min, there are no operations; at 10 min, EBCD is calculated; at 15 min, transport and wet scavenging processes are done; at 20 min, EBCD again; at 25 min, there are again no operations; at 30 min all three processes are calculated, and so on up to 60 min.

In general, when moving from CTM2 to CTM3, there are no changes in the chemistry and aerosol modules, and the CTM2 literature still applies. Some changes are inevitable due to the new core, e.g., the new wet scavenging, and in this section we describe these changes. The main model improvements are listed in Table 1, and will be discussed in these sections: transport (Sect. 2.2), scavenging (Sect. 2.3), calculation of photodissociation rates (Sect. 2.4), emission inventories used (Sect. 2.5) and lightning NO_x (Sect. 2.6). First, we give a short introduction to the chemistry and aerosol schemes.

2.1 Chemistry

The chemistry of CTM3 is identical to CTM2, which has been described in earlier work, comprising comprehensive schemes for both tropospheric and stratospheric chemistry. Tropospheric chemistry was introduced by Berntsen and Isaksen (1997); stratospheric chemistry is based on Stordal et al. (1985), and the tropospheric sulfur chemistry is described by Berglen et al. (2004). The use of full tropospheric and stratospheric chemistry has been reported by, e.g., Gauss (2003), Søvde et al. (2008) and Søvde et al. (2011b). Chemical kinetics are from JPL-06 (Sander et al., 2006) and IUPAC Atkinson et al. (2004).

The tropospheric scheme is a stand-alone module, while the stratospheric module requires tropospheric chemistry to be included. Having the option to turn off stratospheric chemistry is preferable due to computational limits and when the lower troposphere is the domain of interest. In such cases, all the tropospheric trace species are advected throughout the stratosphere but without real chemistry: species that are photochemically destroyed in the stratosphere are allowed to decay at a fixed rate; species with sources in the stratosphere like O_3 and NO_x are set to model climatological values at CTM levels a few km above the tropopause, where the model climatology is produced using CTM3 with full stratospheric chemistry. This will not be described further here.

In this work, we focus on gas phase chemistry. However, as will be described, we also carry out simulations where we include aerosol modules which affect chemistry directly. These are the tropospheric sulfur module comprising sulfur chemistry and sulfate aerosols (Berglen et al., 2004), and nitrate aerosols (Myhre et al., 2006), which affect gaseous HNO_3. The latter also requires sea salt aerosols to be included (Grini et al., 2002). Since the aerosol modules are unchanged from CTM2 to CTM3, we will not evaluate the aerosols here; that is left for a separate study where all available aerosol modules currently in CTM2 (e.g., Grini et al., 2002; Berglen et al., 2004; Grini et al., 2005; Myhre et al., 2006) will be evaluated (in preparation, Søvde et al., 2012). It should be noted that the tropospheric chemistry already includes a parameterization of N_2O_5 conversion to HNO_3 on aerosols, and this treatment is the same in CTM3 and CTM2. The aerosol climatology used is an annual mean height-latitude distribution (Dentener and Crutzen, 1993), and while this is a crude parameterization, an update of this treatment has been outside the scope of this work.

As in CTM2, the chemical integrator is the quasi steady-state approximation (QSSA, Hesstvedt et al., 1978). Tropospheric chemistry is integrated with a maximum time step of 15 min (shorter if the operator-split time step is shorter), except for the OH-chemistry which uses a 1/3 of this time step (Berntsen and Isaksen, 1997). Stratospheric chemistry is carried out with a maximum of 5 min.

2.2 Transport

The model transport covers large-scale advection treated by the second order moments (SOM) scheme (Prather, 1986), convective transport based on Tiedtke (1989) and boundary layer mixing based on Holtslag et al. (1990). The latter is the same in both Oslo CTM3 and Oslo CTM2, and will not be described further.

2.2.1 Advection

In the Oslo CTM3 we have implemented the UCI CTM transport core documented by Prather et al. (2008) (P2008 from now). This new core advection scheme (P2008) is a significant advance in terms of computational efficiency and flexibility compared with the original SOM method used in CTM2 (Prather, 1986).

The 3-D advection is isotropic, i.e. it uses the same SOM algorithm in all dimensions. The zonal (U) and meridional (V) meteorological fields (3-h instant values) are used to compute the vertical (W) field. The CTM keeps and computes only dry-air fluxes, masses and surface pressures. Small inconsistencies in the pressure tendency compared with the convergence fields from $[U]+[V]$ are corrected by small adjustments (typically $<1\%$) in $[U]$ and $[V]$ to ensure consistent mass fields (i.e., the dry-air mass is conserved and never changed except by resolved advective transport). Convection is diagnosed as updrafts and downdrafts separately, taking both entrainment and detrainment into account (see Sect. 2.2.2). The inferred residual subsidence (i.e., a W-like flux) is combined with the large-scale $[W]$ from the $[U]+[V]$ fields to eliminate one advection step.

Advection in each dimension is calculated as a pipe-flow as in P2008, wherein each pipe calculates its own CFL limit (i.e., flux out of a grid box cannot exceed 99 % of the mass) and then if a reduced time step is required, each pipe does multi-stepped advection internally, saving large amounts of computational overhead and not requiring the global calculation to slow down for enhanced multi-stepping in the jet regions. The global advection time step, no longer limited by the CFL criteria, is now limited by the Lifshitz criterion, which is implemented here by requiring that during the sequence of advection, $[\text{Convection}+W]+[U]+[V]$, the mass of any grid box does not fall below 5 % of its initial value. This new CTM3 core advection allows for dynamical time steps as large as 55 min on average for T42L60 ($\sim 2.8°$) fields and 15 min for T319L60 ($\sim 0.55°$) fields, demonstrating the advantage of being Lifshitz-limited for high-resolution CTMs. In contrast to this, the Oslo CTM2 uses the same advection time step at low and high latitudes, and therefore combined high-latitude gridboxes (so-called extended polar zones) during transport, to avoid very short global time steps. Nonetheless, the average CTM2 advection time step for the $[U]$-flux is 10 min, and the averaging

itself makes the treatment less optimal for polar studies, as is shown in P2008 and later in this work.

The over-the-pole flow in P2008 worked well for typical solid-body rotation tests (Williamson et al., 1992), but recently we found that under certain conditions it fails dramatically. First is the relatively large region of high-N_2O air in the Arctic middle stratosphere that splits off from the tropics in March 2005 and survives into August (Manney et al., 2006), a so-called Frozen-In Anti-Cyclone (FrIAC). Second is the Arctic ozone hole of 2011 in which very low O_3 columns are maintained through March and into early April. In both cases an isolated vortex (anti-cyclonic in the first, and cyclonic in the second) has an edge at the pole, and in both cases the P2008 polar advection algorithm failed to correctly transport tracers across the pole, resulting in vortex air being transported out of the vortex and air from outside transported in. With these two test cases, a new over-the-pole treatment was developed.

In the updated, improved SOM version of the P2008 advection algorithm used in CTM3, meridional advection is still calculated as a connected pipe-flow combining a meridion with its antipode complement, but over-the-pole flow at the two pie-shaped grid boxes at each pole is no longer implied. These two grid boxes are no longer combined before the [V] advection step, and all transport in these polar boxes is through the [U]-flux. Given the very small area of these polar-pie boxes, the Lifshitz limiter severely restricted the size of the global time step. Thus, each polar-pie box and its adjacent lower-latitude box are combined (conserving all moments) before both [U] and [V] transport, and restored to individual boxes after, assuming unchanged polar pie air mass (not tracer mass). An alternative version of the SOM polar advection, in which the polar-pie boxes are not combined with their much larger lower-latitude neighbors, is tested here and designated C3_pole. In the optional C3_pole version, the Lifshitz limiter forces global time steps of $\sim 30\,\text{min}$, instead of $60\,\text{min}$, and thus the polar transport is more accurate because of higher spatial resolution at the poles and shorter time steps.

2.2.2 Convection

Convection is calculated using information on convective mass fluxes from the meteorological data, both updrafts and downdrafts. In addition, detrainment and entrainment rates into the updrafts and downdrafts are taken into account. In an updraft (or downdraft), there may be both detrainment and entrainment at the same level, thereby working as a vent for gases as they are transported upwards (downwards). Entrainment (E) and detrainment (D) must balance the convective mass flux, so that for a given layer L we have

$$F_{L+1/2} - F_{L-1/2} = E_L - D_L \qquad (1)$$

where $F_{L+1/2}$ and $F_{L-1/2}$ are convective mass flux through the grid box edges (positive for updrafts and negative for downdrafts). $L+1$ is the level above L. The detrainment rate D_L is positive, and is retrieved from the meteorological data, along with the mass fluxes (F), and from Eq. (1) we calculate E_L. A positive E_L means entrainment occurs, while a negative E_L means additional detrainment is needed to balance the net convective mass flux ($F_{L+1/2} - F_{L-1/2}$).

For the CTM2, no information on detrainment rates is used; only the net entrainment ($E_L - D_L$) is calculated from the mass fluxes. Also, the CTM2 does not use information on downdrafts. In the CTM3, however, the use of downdraft mass flux and detrainment rates may increase the mixing between the convective plume and the surroundings, as explained above. We come back to this in Sect. 3.2.2.

2.3 Scavenging

Scavenging covers dry deposition, i.e. uptake by soil or vegetation at the surface, and washout by convective and large scale rain. Dry deposition rates are unchanged from CTM2 to CTM3 (Wesely, 1989). However, the CTM3 uses a more detailed land use dataset, hence the weighting of deposition rates due to different vegetation categories differs from CTM2. While CTM2 uses only 5 categories in T42 horizontal resolution, the CTM3 uses the $1° \times 1°$ 18-category ISLSCP2 MODIS dataset (9-category ISLSCP88 is also available). Differences in both resolution and vegetation fractions lead to changes up to $100\,\%$ in some grid boxes, while the overall deposition rate pattern is maintained. We treat this change as part of the core update.

We will diagnose wet scavenging more closely in Sect. 3.4, but note here the main differences between CTM2 and CTM3. In general, the ratio of tracer dissolved in rain to that in interstitial air is calculated using Henry's Law. For very soluble species, such as HNO_3, it is often assumed that all is dissolved. When rain containing dissolved species falls from one model gridbox into a gridbox with drier air, it will experience evaporation, depending on the amount of liquid water available (i.e., reversible evaporation). Ice scavenging, however, can be either reversible or irreversible, as explained by Neu and Prather (2012).

The CTM2 large-scale wet scavenging assumes that the gridbox fraction subject to rain (f) is

$$f = c_f \frac{\text{net rain out}}{\text{cloud water} + \text{ice}} \qquad (2)$$

where c_f is cloud fraction and net rain out is the difference between outgoing (grid box bottom) and incoming (top) rain. If net rain out is larger than the available cloud water + ice, CTM2 assumes $f = c_f$. Then the amount dissolved, either from Henry's law or from a fixed fraction, is calculated and transported downwards. Evaporation is based on a similar approach, and occurs (reversibly) only when rain into the gridbox is larger than rain out of the grid box.

CTM3 also uses Henry's Law to calculate the amount of species dissolved in rainfall, but has a more complex cloud

model that accounts for overlapping clouds and rain (Neu and Prather, 2012). This scheme also covers HNO_3 removal on ice for temperatures $258\,K < T < 273\,K$, for which the uptake is calculated by Henry's law modified by a retention coefficient. For temperatures below $258\,K$, HNO_3 uptake on ice follows Kärcher and Voigt (2006) (burial where tracer is uniform through the ice). A discussion on retention coefficients is given in Neu and Prather (2012); conceptually, retention implies that a fraction of the dissolved gas is retained in the ice matrix during freezing of supercooled liquid droplets. Both the retention coefficient formulation and Kärcher and Voigt (2006) burial model assume HNO_3 is incorporated into the ice structure rather than adsorbed on ice surface as in Langmuir uptake (e.g., Tabazadeh et al., 1999). Further details, e.g. on evaporation, which is reversible (but possibly irreversible for ice), can be found in Neu and Prather (2012).

Convective scavenging (Berglen et al., 2004) is unchanged from CTM2 to CTM3, treating convective precipitation as rain. Splitting this precipitation into ice and liquid is beyond the scope of this work. However, due to the new model structure there are some differences in the amounts removed by convective precipitation.

2.4 Photodissociation

For calculation of photodissociation rates, the fast-J2 (Bian and Prather, 2002) in CTM2 has been replaced by fast-JX (Prather, 2009) in CTM3. The differences between fast-J2 and fast-JX are not tested directly here; fast-JX is treated as a part of the model core update. An important update connected to fast-JX (Photocomp2008, 2010) is the new cloud overlap treatment, described by Neu et al. (2007). The only difference from their original treatment is that instead of using multiple cloud profiles in one gridbox column, a single cloud profile is picked randomly from the possible profiles. This will be explained below.

The fast-JX scheme version 6.5 calculates photolysis rates (Js) for stratospheric and tropospheric species from the surface to $60\,km$ for a single column atmosphere defined by CTM layers that can contain both absorbers (O_2, O_3, aerosols) and scatterers (molecules (Rayleigh), aerosols, cloud liquid water, cloud ice water). Based on new data, solar fluxes were revised upward in the 180–$240\,nm$ region and thus photolysis of O_2 increased, resulting in CTM3 showing better agreement with observed O_3 in the stratosphere. Cross sections, quantum yields and solar fluxes are as described by Hsu and Prather (2009), using photochemistry rates from Sander et al. (2006) and cross sections from Atkinson et al. (2004). Photolysis of volatile organic compounds (VOC) is taken solely from the latter, and uses the provided individually tuned Stern-Vollmer pressure dependencies for the quantum yields. With the JPL-2010 update (Sander et al., 2011), most VOCs are now included, and the next release of fast-JX (version 6.7) has a new approach for handling the Stern-

Vollmer pressure dependence of the VOC quantum yields. In version 6.7, the combined cross section plus quantum yield tables for the fast-JX wavelength bins are calculated for three troposphere temperatures with their included pressures for a typical lapse rate: $295\,K$ ($0\,km$, $1000\,hPa$), $272\,K$ ($5\,km$, $566\,hPa$), and $220\,K$ ($13\,km$, $178\,hPa$). Interpolation is done across temperature as before. It should be noted that the CTM2 fast-J2 also uses photochemistry rates from Sander et al. (2006) and cross sections from Atkinson et al. (2004).

The multi-stream scattering calculation assumes horizontally uniform layers. The J-values are generally computed at the mid-point of each layer, with the flux divergence computed between the top and bottom of each layer. When optically thick clouds occur, extra points are inserted within each CTM layer using a logarithmic scale to more accurately (and efficiently) calculate the average J-values throughout the layer.

When cloud fractions are specified for a grid box, an algorithm (Neu et al., 2007) groups the cloudy layers into maximally overlapping connected layers and randomly overlapping groups. This choice can then be used to define a number of independent, single column atmospheres, each with a fractional area. For the T42L60 fields here, the number of single column atmospheres in a box can be quite large, and is truncated to be no more than $20\,000$ by forcing maximal overlapping clouds in the upper layers. In Neu et al. (2007) calculation of average Js over the box is done by quadrature; the fractional areas of nearly clear (cloud optical depth < 0.5), hazy (cirrus), thick (stratus), and opaque (cumulus) are calculated and a sample column atmosphere from each type is selected to represent that region (i.e., Js are weighted by the fractional area of that type). Using the fractional cloud statistics from the T42L60 fields here, the number of calls to fast-JX averages 2.8 per grid box (out of a maximum of 4).

To reduce this computational cost, a new approach is taken in CTM3, where for each grid box column a single column atmosphere is chosen randomly using its fractional area as the likelihood of being selected. The random number generator is always initialized with the same seed value so that simulations are exactly reproducible. For each three-hourly averaged statistics of cloud water and fraction, three randomly sampled single atmospheres are chosen and held fixed for 1 h. This new random cloud-fraction algorithm has been tested against the quadrature method in the UCI CTM and found to produce results (e.g., tropospheric O_3, OH) from five parallel simulations starting with different random seeds that cluster about the quadrature method. All of these results differ clearly from the fast-J2 (CTM2) average cloud fraction approach in which there is never direct sunlight at the surface for boxes with any fractional clouds (see Neu et al., 2007).

In this work, CTM3 uses a climatology for atmospheric black carbon in the fast-JX calculations. Calculated nitrate, sulfate and sea salt aerosols do not affect the calculation of J-values; this will be revised when all CTM2 aerosol modules have been included into the CTM3.

2.5 Emissions

In this work, the model surface emissions are taken from RETRO year 2000 (RETRO emissions, 2006), using diurnal variations for industrial emissions. We have kept the CTM2 treatment of emitting surface NO_x emissions as 96 % NO and 4 % NO_2 to account for unresolved plume processes, for historical reasons based on Petry et al. (1998). Similarly, SO_2 is emitted as 97.5 % SO_2 and 2.5 % sulfate.

Natural emissions are taken from the POET database (Olivier et al., 2003; Granier et al., 2005), while biomass burning is taken from the Global Fires Emission Database version 3 (GFEDv3), scaled to a height distribution available from RETRO. We use monthly mean GFEDv3 data for the years 1997–2010, and we match the emissions to the meteorological year in the model. NO_x emissions are emitted assuming nitrogen content to be 90 % as NO and 10 % as NO_2 (Andreae and Merlet, 2001).

Monthly mean surface mixing ratios are used for CH_4 due to its long lifetime, and will be discussed in Sect. 3.3.2. The mixing ratios are based on CTM2 simulations with CH_4 emissions and surface deposition.

Aircraft emissions are taken from QUANTIFY, for the year 2000 (Owen et al., 2010), and are emitted as NO. No plume processes are parameterized; the CTM2 plume model was left out of this study because it was tailored to a lower resolution and hence no longer appropriate. The old plume parameterization was found to reduce aircraft induced O_3 by ~ 20 % (Kraabøl et al., 2002).

Lightning NO_x emissions are described in Sect. 2.6, and change daily and yearly according to the meteorological year. The climatological mean lightning source amounts to $5\,Tg(N)\,yr^{-1}$.

As the main purpose for this study is to document the Oslo CTM3, we keep the current emission setup with surface and aircraft emissions fixed at 2000 level. Recently more updated datasets have become available, e.g., Lamarque et al. (2010), but we will leave this for later studies.

2.6 Lightning NO_x

Lightning NO_x (L-NO_x) emissions in the model are calculated from the convective fluxes provided by the meteorological input data. In CTM2, L-NO_x emissions were calculated from the product of convective mass flux and cloud-top height (Berntsen and Isaksen, 1999). CTM3 uses an updated algorithm based on cloud-top height (Price and Rind, 1992) with scaling to match lightning flash rates observed by the Optical Transient Detector (OTD, Christian et al., 2003) and Lightning Imaging Sensor (LIS, Christian et al., 1999b). L-NO_x is emitted as NO in both models, and at a given time t the emissions are

$$E(\boldsymbol{x}, L, t) = \bar{f}(\boldsymbol{x}, t)\, \gamma\, v_L \tag{3}$$

where $\bar{f}(\boldsymbol{x}, t)$ is the grid-averaged lightning flash rate at the horizontal location $\boldsymbol{x}$ and at time t, γ is the NO_x yield per flash, and v_L is the fraction of total column emissions occurring at level L (see below). We choose $\gamma = 246\,mol\,flash^{-1}$, so that the mean climatological emissions are $5\,Tg(N)\,yr^{-1}$.

Lightning is diagnosed in the Oslo CTM3 only when convection out of the boundary layer exceeds a grid-averaged updraft velocity of $0.01\,ms^{-1}$ at $\sim 850\,hPa$, as specified in the input meteorology. In addition, the surface must be warmer than 273 K and the cloud top colder than 233 K to support charge separation within the cloud (Williams, 1985). When these conditions are met, the in-cloud flash rate (f) has the same functional dependence on cloud-top height $H(\boldsymbol{x}, t)$ as in Price and Rind (1992):

$$f(\boldsymbol{x}, t) = \begin{cases} \alpha_l H(\boldsymbol{x}, t)^{4.9} & \text{over land} \\ \alpha_o H(\boldsymbol{x}, t)^{1.73} & \text{over ocean} \end{cases} \tag{4}$$

Cloud-top heights in each model column are defined as the top of the highest grid box with positive upward convective mass flux. We develop new scale factors, $\alpha_l = 3.44 \times 10^{-6}\,s^{-1}\,km^{-4.9}$ and $\alpha_o = 2.24 \times 10^{-3}\,s^{-1}\,km^{-1.73}$, to match the climatological flash rates over land and ocean, $35\,s^{-1}$ and $11\,s^{-1}$, respectively. The factors are calculated for our meteorological data over 1999–2009 (ECMWF cycle 36r1, resolution T42L60), constrained by OTD-LIS observations during 1995–2005. These coefficients differ from those reported by Price and Rind (1992) by up to a factor of 10, and must be re-calibrated for other meteorological data, and also for different resolutions, e.g. T319.

Convective fluxes, and the clouds that contain them, occupy a small fraction of the horizontal area of a model grid square. Therefore, lightning also occupies a small fraction of the model grid area. The grid-averaged lightning flash rate ($\bar{f}$) is then

$$\bar{f}(\boldsymbol{x}, t) = a(\boldsymbol{x}, t)\, f(\boldsymbol{x}, t). \tag{5}$$

where a, the area fraction experiencing lightning, accounts for the variation in grid size with latitude and is estimated as a weighted average of the cloud fractions in the column:

$$a(\boldsymbol{x}, t) = \frac{\sum_{l=1}^{l_t} f_{c,l} w_l}{\sum_{l=1}^{l_t} w_l}. \tag{6}$$

Here, $f_{c,l}$ is the ECMWF cloudy fraction at level l (McCarty et al., 2012), w_l is the wet convective mass flux at level l, and l_t is the model level at the cloud top.

Figure 1 compares the lightning distributions in CTM2 and CTM3 with OTD-LIS observations. CTM3 has much less lightning over oceans than CTM2, especially around Indonesia, but still more than the observations. In both models, flash rates are underestimated in Africa and overestimated in South America, but these biases are slightly reduced in CTM3. The convective parameterization used in
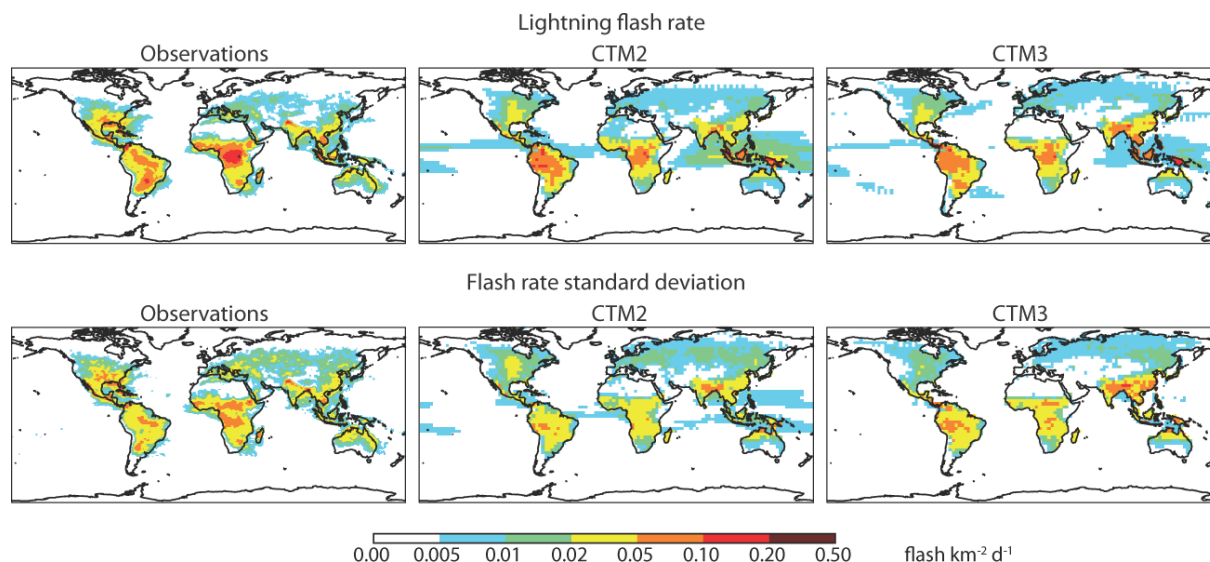
Fig. 1. Climatology of lightning flashes in observations and models. Annual mean flash rate (top) and standard deviation of monthly flash rates (bottom). Observation panels show OTD-LIS during 1995–2005. Model panels show Oslo CTM2 and Oslo CTM3 for 1999–2009.

ECMWF meteorology may contribute to these lightning biases (Barret et al., 2010), although many models with different convective schemes exhibit similar problems over these continents (Labrador et al., 2005; Tost et al., 2007). This feature could be a weakness of the Price and Rind (1992) equations; Barret et al. (2010) reported similar behavior for several models using the Price and Rind equations. CTM3 also better reproduces the seasonal north–south migration of lightning in Africa, as seen in the monthly standard deviations (Fig. 1, bottom row). Overall, the spatial distribution of L-NOx emissions in CTM2 and CTM3 are similar, but CTM3 better matches the OTD-LIS observations.

The model distributes L-NO_x emissions vertically through the convective column according to observed profiles (Ott et al., 2010) for 4 world regions. These profiles are scaled vertically to match the height of each convective plume in the CTM and already account for vertical mixing of lightning NO_x within the cloud. Geographic region definitions are from Allen et al. (2010) and Murray et al. (2012).

The old vertical profiles (Pickering et al., 1998) injected a large fraction of L-NO_x near the surface and near the convective cloud top. In contrast, the new vertical profiles (Ott et al., 2010), which are based on more extensive in situ measurements, place most L-NO_x below the convective cloud top with little near the surface. While the new L-NO_x algorithm scales the new profiles to match the cloud-top heights hourly in each CTM column, the old CTM2 algorithm assumed a fixed convective top at 16 km for purposes of calculating the vertical distribution of L-NO_x. Thus, CTM2 placed much more L-NO_x in near the tropopause, or even in the stratosphere when the tropopause was below 16 km. The change in L-NO_x between CTM3 and CTM2 is extensive, includ-

ing location, flash-rate and scaling factors as well as vertical profiles of the injected NO.

We have tested the change in profiles in CTM3, and found that compared to the old profiles (Pickering et al., 1998), the Ott et al. (2010) profiles cause modelled zonal mean NO_x to increase by up to 10–15 % (annually up to ~ 10 %) in the middle troposphere (400–800 hPa), and decrease NO_x by up to 15–25 % (~ 15 % annually) in the tropical upper troposphere (250–100 hPa). Accompanying O_3 changes range from -2 % to 2 %. However, scaling the old profiles to fixed 16 km convective cloud tops has a larger effect, increasing NO_x by more than 100 % at ~ 200 hPa and O_3 by ~ 35 % at 400 hPa. We come back to this in Sect. 3.3.

3 Evaluation

Here we evaluate the Oslo CTM3 against the Oslo CTM2 and measurements. The comparison is carried out in several steps, focusing first on the stratosphere and then on the troposphere, before looking at a number of diagnostics. For this study, we have carried out several full chemistry-transport simulations, listed in Table 2. Also listed are the time spans simulated for the different runs. CTM3 and CTM2 are driven by the same meteorological dataset.

The main runs are C3, C3_ssn, C2 and C2_ssn, and the others are sensitivity studies. The ssn-simulations include sulfur chemistry (Berglen et al., 2004), sea salt aerosols (Grini et al., 2002) and nitrate aerosols (Myhre et al., 2006). As will be explained, the C3 simulation produces higher tropospheric OH and O_3 than C2, due to less removal by wet scavenging (e.g., HNO_3 and H_2O_2), different vertical distribution of lightning, and more active photochemistry. This

Table 2. Description of the full transport-chemistry simulations in this study. The notation C2 and C3 represents CTM2 and CTM3, respectively. Lightning parameterizations are described in Sect. 2.6 The $[U]$-wind time step annual average for year 2005. "SSN" denotes "sea salt, sulfate and nitrate". * C3_1/4 was started from C3_1/2 at 1 Jan 2005.

Simulation	Start	End	Lightning (see text)	Detrainment rates used?	Chemistry Trop/Strat/SSN	Operator-split time step	$[U]$-flux time step
C3	1 Jan 1997	31 Dec 2005	OTT	Yes	Y/Y/N	60 min	55 min
C3_ssn	1 Jan 1997	31 Dec 2005	OTT	Yes	Y/Y/Y	60 min	55 min
C2	1 Jan 1997	31 Dec 2005	C2PIC	no	Y/Y/N	60 min	10 min
C2_ssn	1 Jan 1997	31 Dec 2005	C2PIC	no	Y/Y/Y	60 min	10 min
C3_1/2	1 Nov 2004	31 Dec 2005	OTT	Yes	Y/Y/N	30 min	30 min
C3_1/4	1 Jan 2005*	31 Dec 2005	OTT	Yes	Y/Y/N	15 min	15 min
C3_pole	1 Nov 2004	31 Dec 2005	OTT	Yes	Y/Y/N	60 min	29 min

motivated the inclusion of the ssn-studies, which were known to reduce OH. Comparisons with observations will in general be made with C3_ssn and C2_ssn, although we will also comment on C3 and C2 when these differ from the ssn-results.

The sensitivity studies are all without sulfur and nitrate aerosols, included to assess the CTM3 transport, and are compared with C3. C3_1/2 is the C3 run with an operator-split time step of 30 min, i.e. half of that in C3. It is included to assess transport errors. Similarly, we carry out a 12 month C3_1/4, starting at 1 January 2005 from the C3_1/2 simulation, for testing the convergence of operator-split time steps. C3_pole is the C3 run with the optional more accurate polar cap treatment described in Sect. 2. As mentioned in Sect. 2, CTM3 transport is stable for longer time steps than CTM2, and may produce larger errors e.g. around jet streams. It is the $[U]$-flux time step that largely controls these errors, and from Table 2 we also show how these differ for the different CTM3 setups.

For the main runs, the simulated period is 1997–2005, while the other runs are 14-months simulations starting from the instant C3 field at 1 November 2004 (except the 12-month C3_1/4, started from C3_1/2 at 1 Jan 2005). The first year (1997) was started from an already spun-up simulation, more specifically from an instant snapshot of all species at 1 January 2005, and should therefore be considered as spin-up. The year 2005 was chosen because the Northern Hemisphere meteorological conditions were somewhat similar to 1997.

3.1 Stratosphere

For the evaluation of the stratosphere, we first consider the stratospheric age of air and how the age tracers differ from CTM2 to CTM3. In addition, we carry out comparisons with satellite measurements. It should be noted that in the stratosphere, C3_ssn and C3 are only negligibly different because their main differences are in the troposphere.

3.1.1 Age of air

The UCI CTM/Oslo CTM3 stratospheric age of air was presented by Prather et al. (2011), using an older cycle of the ECMWF IFS meteorology, and for a different year. In the tests here with CTM2 and CTM3, the purpose of examining age of air is to identify differences in the tracer transport algorithm and the impact of polar treatment and time step. All versions use the identical three-hour meteorological data. Our intent is not to evaluate these met fields against observations of tracers that approximate the age of air (e.g., SF_6, CO_2), but to examine how different numerics can produce different values (e.g., P2008). We calculate stratospheric age of air from a tracer that is forced to be linearly increasing in the lower tropical troposphere (Hall et al., 1999). The age is calculated for T42 horizontal resolution, and the transport runs are similar to the runs listed in Table 2.

Simulations of 20 years were carried out using 2005 meteorology recycled annually, and age of air was calculated from the linearly increasing source in the tropical troposphere. Note that the use of a tropical source gives greater north–south symmetry in the lower stratosphere than from northernly-only sources used in observations, see Prather et al. (2011). Annual zonal means of age of air are shown in Fig. 2 for CTM3 and CTM2, with operator-split time step of one hour (C3 and C2), along with C3_1/2 where the time step is 30 min. In general, the maximum age value is slightly smaller in CTM3. However, the main difference between CTM3 and CTM2 can be found in the Southern Hemisphere for ages older than 4 yr, where the CTM2 produce a southward tongue of younger air at around 40° S–70° S and 50 hPa–10 hPa altitude. This feature is not captured by C3, but is better captured in C3_1/2, where the shorter operator-split time step reduces the errors, especially around the polar jets (Prather et al., 2008). Between C3 and C2, the maximum age difference in this region amounts to 0.67 yr, as shown in Fig. 3a. By halving the operator-split time step as in C3_1/2, the maximum difference to C2 at this location is reduced to 0.29 yr (Fig. 3b). An operator-split time step of 15 min reduces the difference to C2 further to 0.12 yr (Fig. 3c). Thus, the error induced by a finite time step can be readily evaluated from the sequence C3, C3_1/2, C3_1/4, and it is not clear whether the differences in Fig. 3c are from errors in C3_1/4 or C2.
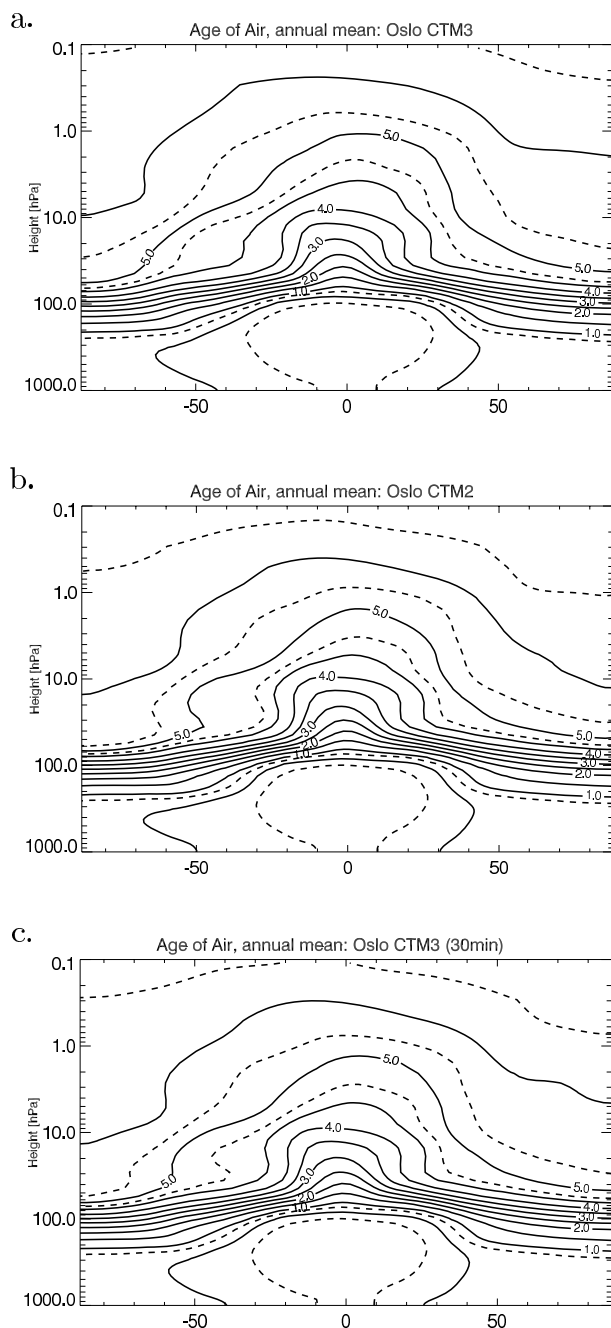
Fig. 2. Annual zonal mean age of air based on a tropical source tracer increasing linearly with time, after 20 yr of transport, resolution T42L60. **(a)** CTM3, **(b)** CTM2 and **(c)** CTM3 with 30 min operator-split time step.

Fig. 3. Difference in annual zonal mean age of air between C3 and C2 **(a)**, C3_1/2 and C2 **(b)** and C3_1/4 and C2 **(c)**.

Whereas a comparison of C3_1/2 against C3 shows a noticeable improvement when reducing the time step, the improved polar cap transport (C3_pole) only changes the stratospheric age of air slightly (not shown). From a latitude/longitude view, the difference between CTM2 and CTM3 at ∼ 50 hPa (not shown) is found around the polar vortex edge, indicating that the vortices are more closed off in CTM2. Halving the time step removes most of this dissimilarity (Prather et al., 2008), consistent with the better resolved tongue of younger air seen in Fig. 2c.

This indicates that CTM3 should be run with a 15 min or 30 min operator split time-step in order to capture the correct distribution and transport at the polar vortices. We come back to this in Sect. 3.5.
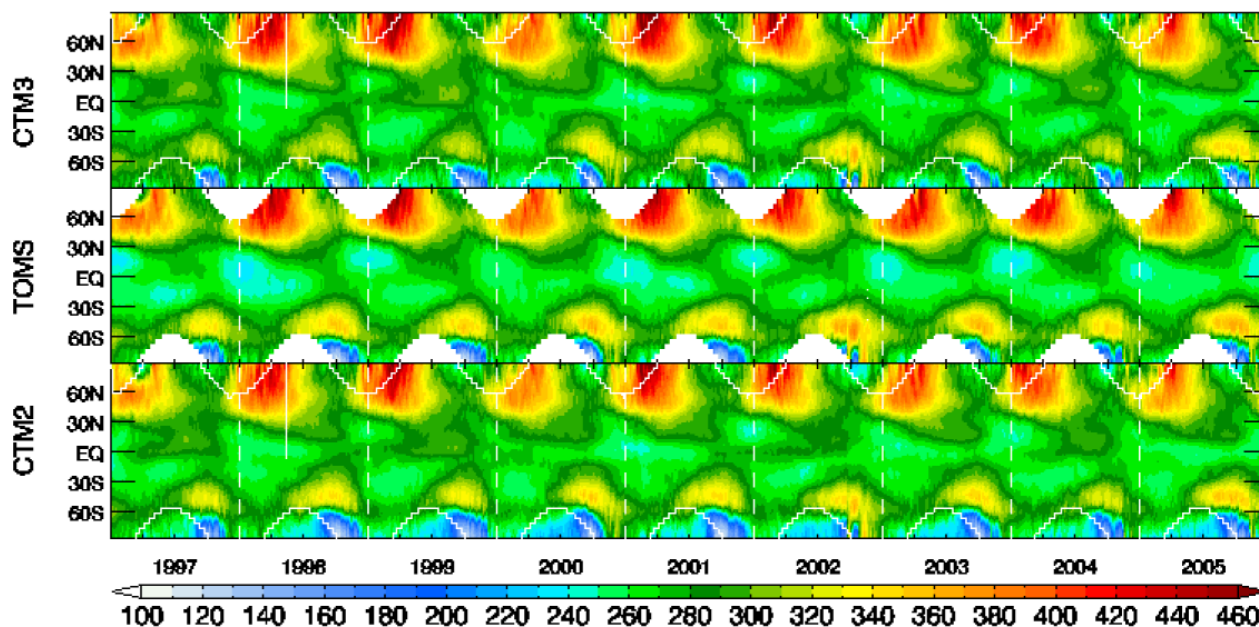
Fig. 4. Daily zonal mean total O_3 column (Dobson Units) for C3_ssn (top), TOMS (middle) and C2_ssn (bottom), from 1997 through 2005.

3.1.2 Comparisons with satellites

In Fig. 4 we show the interannual variability of the daily zonal mean O_3 column, from 1997 to 2005, where model results from CTM3 (C3_ssn) and CTM2 (C2_ssn) are compared to Total Ozone Mapping Spectrometer (TOMS) instrument, available from http://mirador.gsfc.nasa.gov/. Figure 4 shows clearly that the models produce reasonable polar O_3 loss, as previously reported (Søvde et al., 2008, 2011b; Balis et al., 2011). Other features that are driven by the meteorological dynamics and the seasonality of the amount of sunlight available can also be seen, such as the Antarctic Ozone Hole split in 2002.

By closer examination, we see that the Oslo CTM3 and CTM2 O_3 columns in general agree well with TOMS in the Northern Hemisphere (NH). At Southern Hemisphere (SH) mid-latitudes, CTM3 is somewhat lower than observed and CTM2 seem to reproduce TOMS better, confirming that CTM2 has a somewhat more closed-off polar vortex, and that the CTM3 transport errors are larger near the SH polar jet. In 2002, however, when the Antarctic O_3 hole was smaller and also split in two, the models are more similar at SH mid-latitudes, indicating that the over-shooting winds were less pronounced. In Fig. 5 we show the CTM3 vs. CTM2 difference for year 2005, for different operator-split time steps in CTM3. Clearly, CTM3 converges towards CTM2 for shorter time steps, with the 15 min time step close to the ~ 10 min [U]-flux time step in CTM2. We note, however, that CTM2 does not necessarily show the truth and also has its errors, especially near the pole where CTM3 resolves transport better.

The use of real meteorological data allows a more thorough comparison on a one-to-one basis. As an example,

we show in Fig. 6 the daily average total O_3 column for 28 November 2000, when an O_3 mini-hole occurred over Europe (Semane et al., 2002). Both CTM3 and CTM2 reproduce this event well and to a better degree than in Søvde et al. (2008), which had too high O_3 columns for Northern Europe towards Siberia. This improvement is mainly due better stratospheric circulation in the 60-layer meteorological data used here compared to the 40-layer data used in Søvde et al. (2008).

It can be seen in Fig. 6 that the CTM2 produces almost uniform polar caps due to its transport treatment. CTM3, however, does not, and reproduces, e.g., the Antarctic low-O_3 tongue at about 170° E to a better degree. In agreement with Fig. 4, the SH mid-latitude maximum is somewhat higher in CTM2. Similar features can be found in 2005, where we have found that this bias is reduced in C3_1/2.

The Oslo CTM3 mostly produces higher column values than CTM2, the latter having a slightly lower stratospheric O_3. This is shown in Fig. 7, where the models are compared to Earth Observing System Microwave Limb Sounder (MLS) version 3.3 observational data (Froidevaux et al., 2008; Livesey et al., 2011). Zonal means for 30-degree latitude bands (60° for the tropics) are given for January 2005 and July 2005, with standard deviation shown as horizontal bars. CTM2 standard deviation is not shown; it is similar to CTM3. For O_3 (Fig. 7a), CTM3 reproduces MLS better than CTM2 due to the updated photolysis treatment. The sensitivity studies C3_1/2 and C3_pole produce negligibly different profiles (not shown).

In Fig. 7b we compare the models with MLS HNO_3 measurements (Santee et al., 2007; Livesey et al., 2011), where
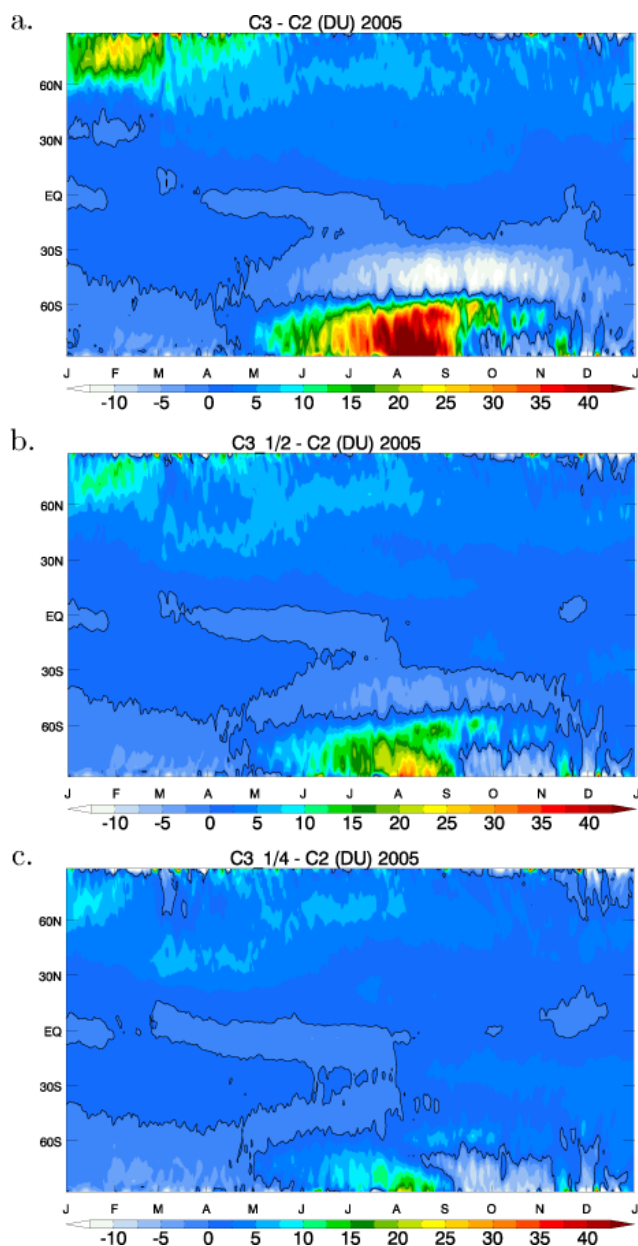
Fig. 5. Convergence of the CTM3 operator-split time step, shown by the difference in daily zonal mean total O_3 column (Dobson Units) between CTM3 runs with different operator-split time steps and CTM2, for the year 2005. **(a)** C3 (60 min) vs. C2 **(b)** C3_1/2 (30 min) vs. C2 and **(c)** C3_1/4 (15 min) vs. C2.

the CTM3 and CTM2 reproduce MLS well in summer but underestimate at altitudes above 30 hPa in winter. The CTMs already include NO_x conversion to HNO_3 on background aerosols (Søvde et al., 2008), so an underestimation of HNO_3 in winter may, e.g., be due to lack of nitrogen species transported downwards from the mesosphere (Randall et al., 2006, 2009), or the lack of in-situ NO_x sources caused by energetic particle precipitation (Jackman et al., 2008; Semeniuk et al., 2011) and conversion to HNO_3 by, e.g., ion clusters (Verro-

nen et al., 2008, 2011). Below 30 hPa the models do fairly well, also when it comes to the standard deviation. There are no big differences between CTM2 and CTM3, although CTM3 seems to perform slightly better in summer. Again, C3_pole and C3_1/2 produce almost identical profiles, except the latter at SH high latitudes, where small differences up to 5 % in HNO_3 can be seen (not shown).

Lastly, we compare modelled N_2O with MLS measurements (Lambert et al., 2007; Livesey et al., 2011), as shown in Fig. 7c. CTM3 and CTM2 produce very similar distributions of N_2O. However, at all latitudes except high wintertime latitudes, both models underestimate N_2O between about 30 hPa and 1 hPa, indicating slower Brewer-Dobson circulation in the meteorological data, particularly for upward motion in the tropics. Again there are negligible differences to C3_1/2 and C3_pole. An earlier cycle of meteorological data (cycle 29) shows slightly better comparison (not shown), which could indicate that cycle 36r1 has slower vertical transport in the tropics. Monge-Sanz et al. (2012) also found slower Brewer-Dobson circulation in the ERA-Interim forecasts (cycle 31r2), producing too old stratospheric air in the tropics at 18–22 km altitude. They found that the reanalysis produces better age of air in the tropics, but that forecasts are better at other latitudes. We will come back to the slow Brewer-Dobson circulation in the next sections.

3.1.3 Modelling a Frozen-in Anti-Cyclone

An important change from CTM2 to CTM3 is the better polar transport treatment, and to demonstrate this we look at the 2005 Frozen-in Anti-Cyclone (FrIAC) in the Arctic reported by Manney et al. (2006) and more recently studied by Allen et al. (2011). In Fig. 8 we show a Hovmöller plot, as in Allen et al. (2011), of N_2O at the altitude 850 K and latitude 78° N, measured by MLS, and modelled by the Oslo CTM3 and the Oslo CTM2. MLS measurements have been binned into $5 \times 5°$ latitude/longitude bins due to the sparsity of observations. N_2O from the CTMs were put out hourly in 3-D and interpolated to 78° N. Note that the color scale range of 0–175 ppb is larger than in Allen et al. (2011). We have added contours of 75 ppb and 100 ppb to make a comparison easier.

The Oslo CTM3 reproduces transport of the high-N_2O intrusion well, although the mixing ratio amplitude is underestimated by 40–60 ppb. As shown in Fig. 7, N_2O was in general underestimated between 30 hPa and 1 hPa at most latitudes, which could explain the overall low N_2O abundances relative to observations in Fig. 8. This could render N_2O values too low even before entering the polar latitudes. The Oslo CTM2 does not capture this transport well. Its transport pattern is similar to CTM3, but due to the averaging of the polar gridboxes, where the N_2O is smeared out, the values are substantially lower than for CTM3.

From sensitivity studies, we have found (but not shown) that halving the operator-split time step (C3_1/2) does not change the CTM3 performance in terms of the N_2O levels
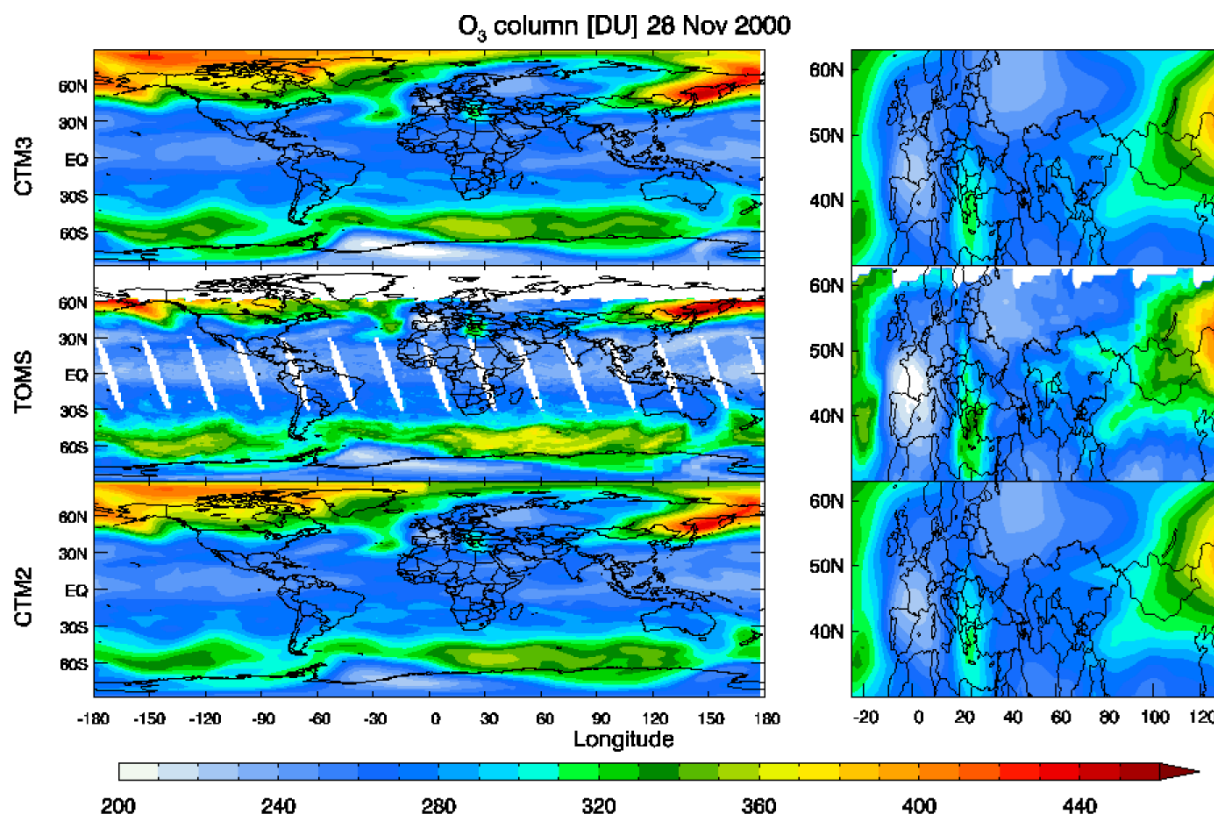
Fig. 6. Daily mean total O_3 column for C3_ssn (top), TOMS (middle) and C2_ssn (bottom), 28 November 2000.

in the FrIAC. This is mainly because the FrIAC was located above the region where main transport differences between C3 and C3_1/2 are found. However, the improved polar transport (C3_pole) produces up to 30 % more N_2O in the FrIAC after 1 May. In general, the differences arise when large wind gradients are located across the combined grid boxes in CTM3 transport.

Studies involving polar cap transport, and to some extent Arctic and Antarctic studies, will clearly benefit from using C3_pole transport, while other studies may benefit from the shorter computing time achieved by combining polar boxes, and find its performance acceptable.

3.2 Troposphere

In this section we focus on the troposphere, describing transport differences between Oslo CTM3 and CTM2 and also compare the models with observations.

3.2.1 Transport in the troposphere

CTM3 and CTM2 are expected to differ slightly in the troposphere due to the differences in transport treatment. We have studied the linearly increasing age of air tracer for year 20, and in the zonal mean C3 differs from C2 by only ~ 0.5 % between the surface and ~ 300 hPa (not shown). In general the

CTM3 tropospheric mixing ratios are slightly higher south of $50°$ S, indicating slightly faster transport to southernmost latitudes. North of $50°$ S the differences are generally related to convective regions, where CTM2 has higher mixing ratio than CTM3 in the lower troposphere. This is mainly due to differences in entrainment and detrainment, and will be explained in Sect. 3.2.2.

3.2.2 Convective transport

Recently, Hoyle et al. (2011) presented convective tracer transport using different CTMs, among them the Oslo CTM2, using a different set of meteorological data than used here. To compare the differences in transport between Oslo CTM3 and CTM2, we do similar tracer studies with one tracer held constant at 1 ppm below 500 m altitude and above having a lifetime of 6 h (T6h), and a second 20-days lifetime tracer held constant at 1 ppt at the surface (T20d).

Both models have an overall operator-split time step of 60 min, where large-scale advection is treated for 60 min and then boundary layer mixing for 4×15 min. Constant T6h and T20d values (as described above) are set before every 15-min boundary layer mixing step, so that each mixing step will not experience a surface layer with almost no T6h or T20d tracer available. This is mainly important for T20d, which is only set in the surface layer. The T6h and T20d results
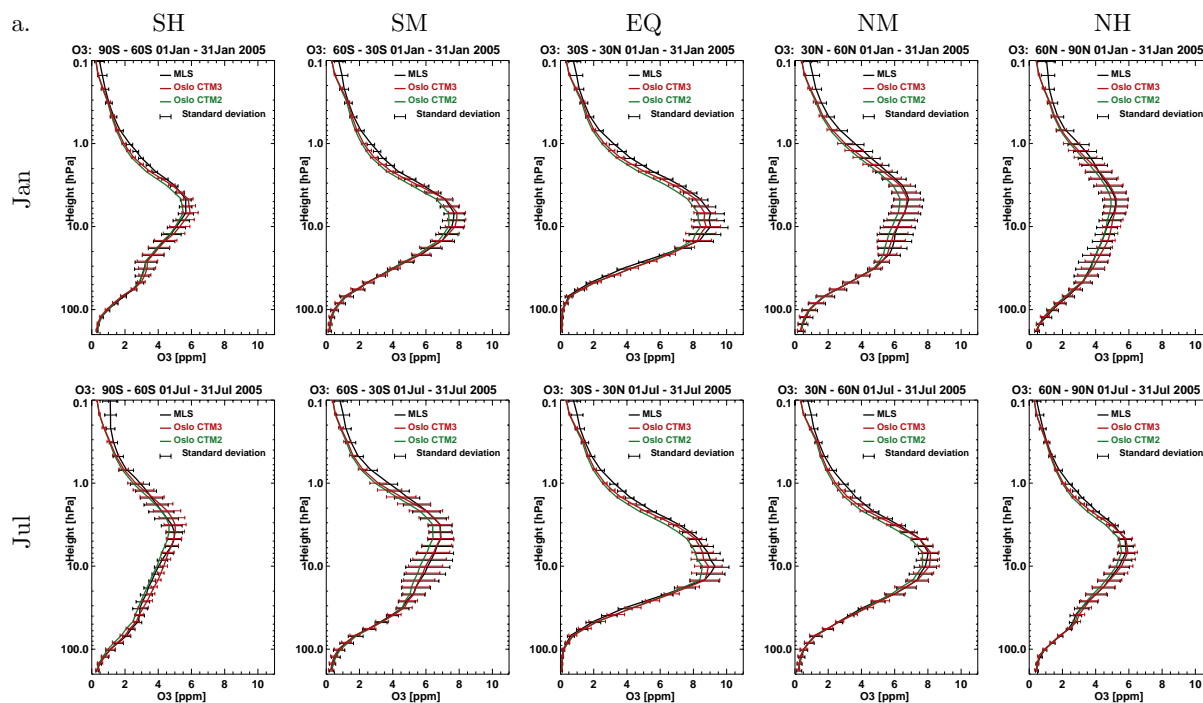
a.

Fig. 7. Comparison of O_3 (**a**), HNO_3 (**b**), and N_2O (**c**) from CTM3 (C3_ssn), CTM2 (C2_ssn) and MLS, as vertical profiles of zonal monthly means covering the latitude bands 90° S–60° S (southern high, column "SH"), 60° S–30° S (southern mid-lat, column "SM"), 30° S–30° N ("EQ"), 30° N–60° N ("NM") and 60° N–90° N ("NH") for January 2005 and July 2005. Model profiles are processed with the MLS averaging kernel and a-priori profiles.
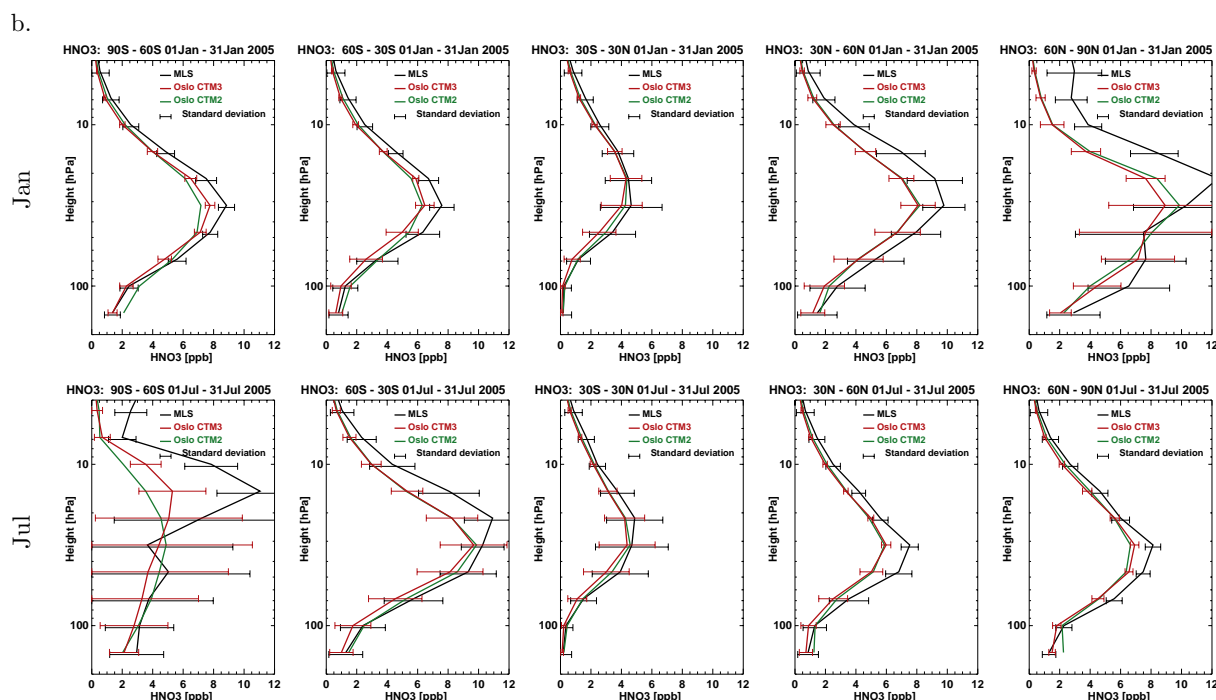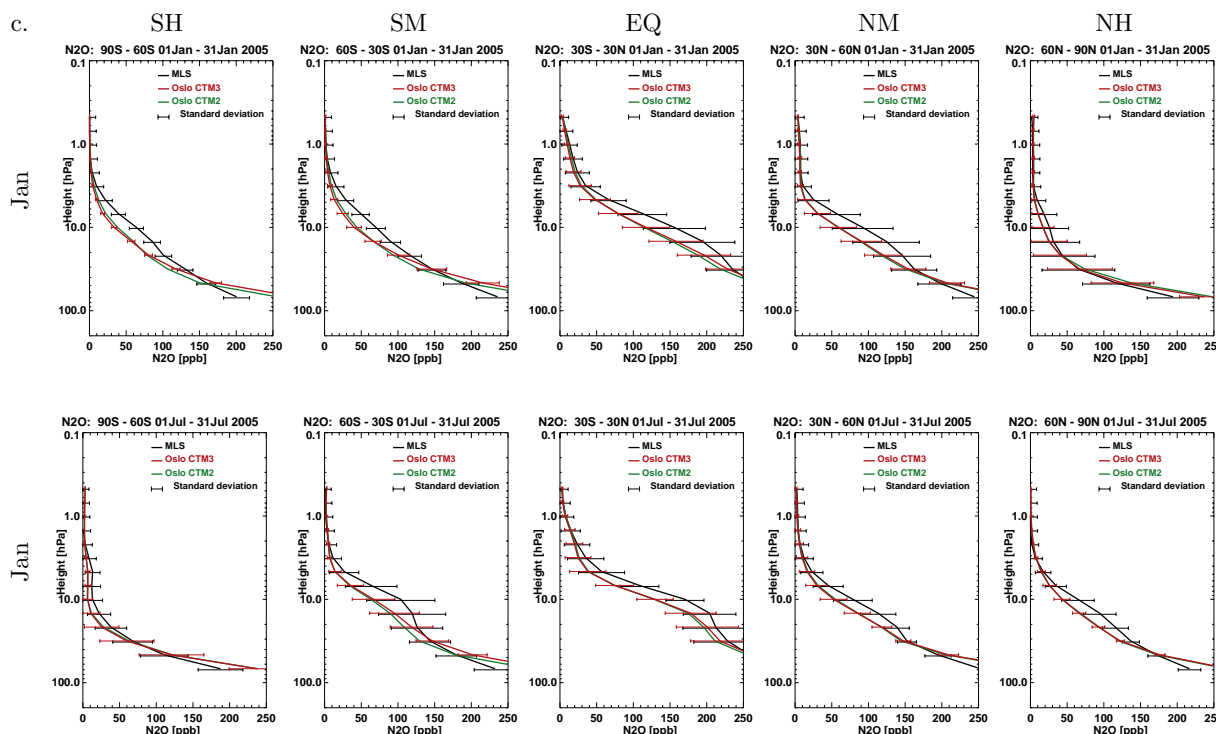
b.

Fig. 7. Continued.

Fig. 7. Continued.

for the same regions as in Hoyle et al. (2011), are shown in Fig. 9, where Oslo CTM3 in general has smaller mixing ratios for T6h than does the CTM2 (up to 100 % difference at 200–100 hPa), and slightly smaller for T20d.

As described in Sect. 2.2.2, the CTM2 only uses updraft convective mass fluxes; convective downdrafts are not taken into account, and neither are detrainment rates into the up- or downdrafts. In Oslo CTM3, downdrafts and detrainment rates are also taken into account, and while downdrafts change the results negligibly (not shown), the use of detrainment rates explains most of the model differences in T6h above 550–300 hPa (dash-dotted lines in Fig. 9). At the surface there are little differences in T6h between CTM3 and CTM2 due to the constantly replenishing to 1 ppm below 500 m. Using the detrainment rates transports substantially less to high altitudes due to venting below, as explained in Sect. 2.2.2. Even though the difference between CTM2 and CTM3 is smaller for T20d than for T6h, skipping downdrafts and detrainments still shifts T20d in CTM3 towards the CTM2 profiles above 700–550 hPa.

Hoyle et al. (2011) used a different set of meteorological data, so a direct comparison is not possible. However, from our results we can assume that inclusion of downdrafts and detrainments would shift their Oslo CTM2 towards the FRSGC/UCI results for T6h below 100 hPa, while T20d would be shifted closer to the other models.

In general, the Oslo CTM3 transports up to ∼ 10 % less out of the lowermost model layers than CTM2, mainly in sub-tropical and mid-latitude regions. This may be due to small differences in the boundary layer schemes or in the different treatments of convection.

3.2.3 Vertical profiles – O₃ sondes

Vertical profiles of O_3, interpolated linearly to the location of selected sonde stations around the world, are put out hourly from the models. This allows reasonable temporal interpolations to sonde launch times, thereby giving a better basis for comparing modelled and observed profiles. To evaluate the modelled O_3 in the troposphere, we compare the models to O_3 sonde measurements available from the World Ozone and Ultraviolet Radiation Data Centre (WOUDC), and also from Southern Hemisphere ADditional OZonesondes (SHADOZ) (Thompson et al., 2003), for the year 2005.

In Fig. 10 we show sonde comparisons for selected stations as monthly means, using model profiles only at measurement times. Observations are in black, CTM3 in red and CTM2 in green. To calculate the means, all profiles for a given month and station have been interpolated to a fixed pressure spacing, i.e. the 60-layer model spacing for a surface pressure of 1000 hPa. For each mean profile we show the standard deviation as horizontal bars, and the range of O_3 in the observations (backslashed black) and in the Oslo CTM3 profiles (slashed red). We show only the mean for CTM2, as its variation is similar to CTM3. In general, CTM3 produce somewhat better profiles than does
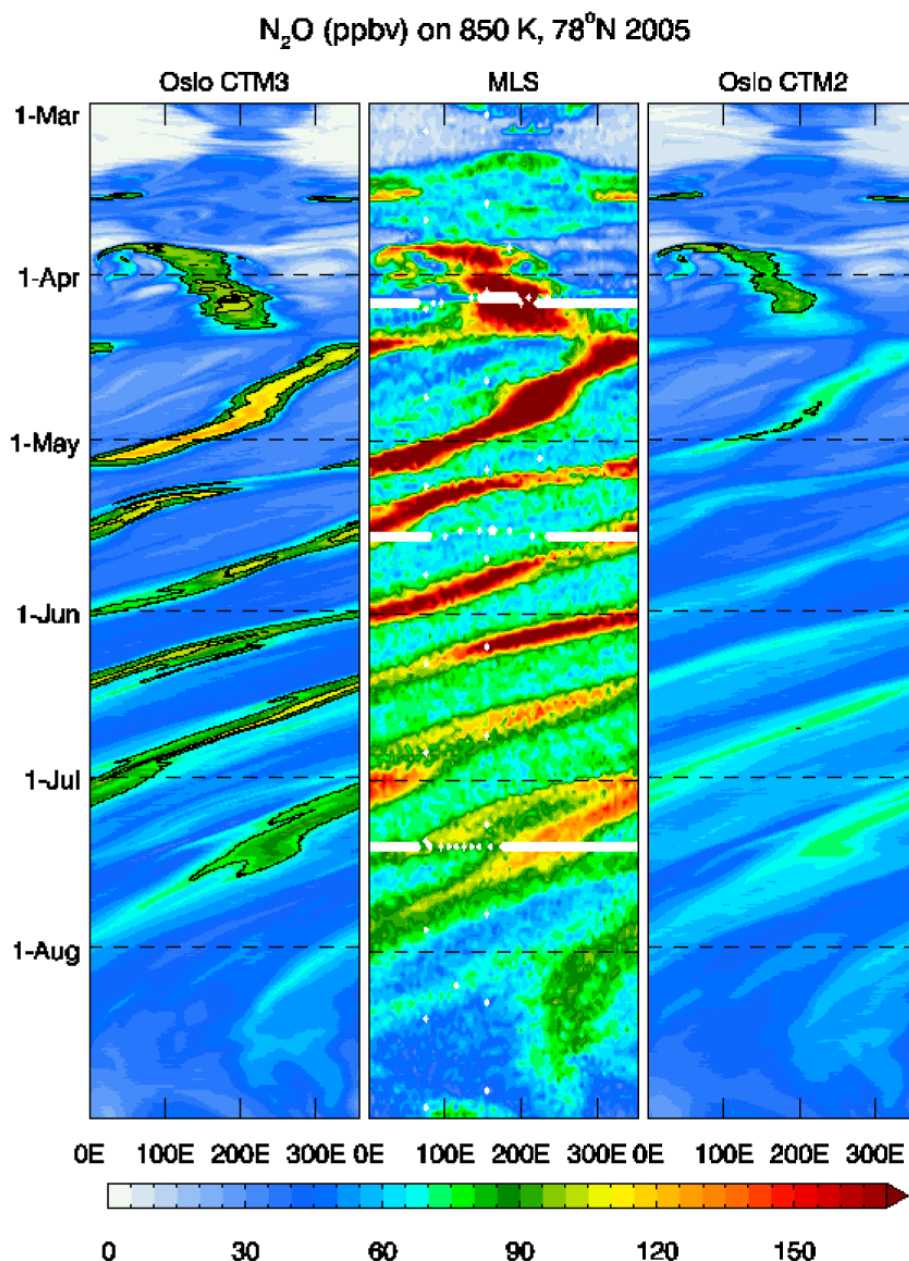
N$_2$O (ppbv) on 850 K, 78°N 2005

Fig. 8. Hovmöller plot of N$_2$O at 850 K and 78° N, as modelled by Oslo CTM3 (left), measured by MLS (middle) and modelled by Oslo CTM2 (right). Contours on CTM panels are 75 ppb and 100 ppb.

CTM2. In the supplementary material we also include one-to-one comparisons of models and observations for all the single sonde measurements used in the means. From these single sonde profiles, we see that CTM2 sometimes capture tropopause folds better than CTM3. However, shortening the transport time step does improve CTM3 tropopause folds (not shown).

As noted in Sect. 2.5, model emissions are for the year 2000. Therefore we have included a similar comparison for the year 2000 in the Supplement, showing similar model performance as for 2005.

In general, CTM3 produce slightly better profiles than does CTM2, although for some of the locations and months, CTM2 reproduce measurements better. There are small differences between C3_ssn and C3, and between C2_ssn and C2, mainly in the upper troposphere and lowermost stratosphere; however, the relative differences near the surface may be larger (up to about 25 %, not shown).
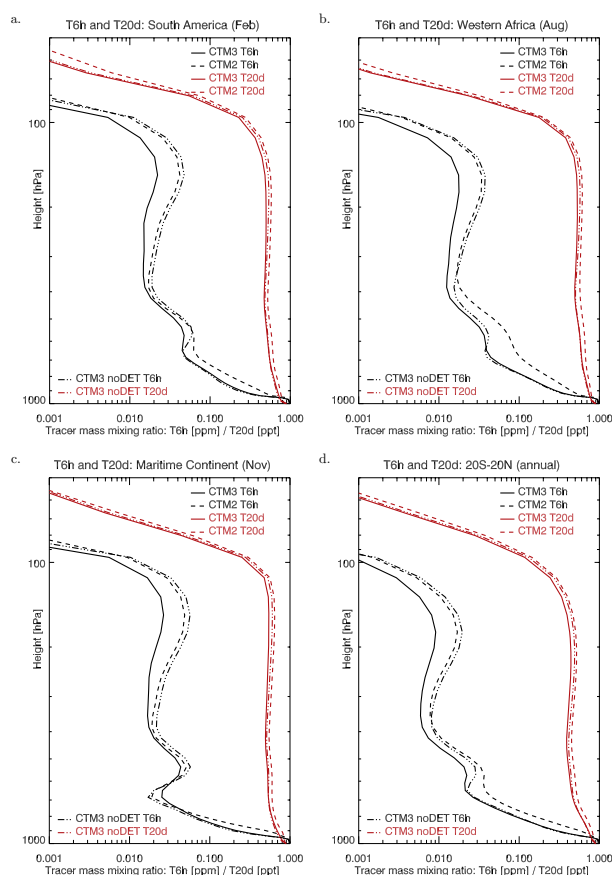
Fig. 9. T6h (ppm) and T20d (ppt) tracer as described in text and by Hoyle et al. (2011): selected monthly means for regions and annual mean for 20° S–20° N.

3.2.4 Modelling CARIBIC CO measurements

The atmospheric abundance of CO is important for OH. On a global average, we find surface CO in CTM3 to match CTM2. On a monthly basis, CTM3 has a more pronounced seasonal variation, being up to 5 % lower in NH spring and 5 % higher in NH autumn. As found in Sect. 3.2.2, CTM3 transports less out of the boundary layer, and for CO we see that CTM3 has up to 50 % higher abundances over land, while having up to 10 % less over ocean. A comparison of modelled and measured CO at Alert station (shown in Supplement) reveals that while CTM3 is 5–15 % lower than CTM2, both models catch the seasonal variation, but underestimates at NH winter, as was found in the multi-model CO-comparison of Shindell et al. (2006). A comparison for the Hohenpeißenberg station, shows that CTM3 is on average about 15 % higher than CTM2. Unfortunately, it has not been possible to compare our results directly with CTM2 results from the comparison of Shindell et al. (2006), a study which used different emissions; with slightly higher total amount of CO emissions, 1077 Tg yr^{-1} compared to our 1012 Tg yr^{-1} from RETRO, POET and GFEDv3 datasets.

It can be noted that the anthropogenic emissions in RETRO do have a month-to-month variation, in contrast to the emissions in their study. Shindell et al. (2006) suggest NH emissions are too low, and recently other studies also suggest that CO emissions are low in general (Lamarque et al., 2010; Pison et al., 2009; Kopacz et al., 2010). It should be mentioned that the recent JPL recommendations (Sander et al., 2011) may change the modelled CO, a study which has not been possible in this work.

With our focus on transport differences between CTM2 and CTM3, we look more specifically at the capability of CTM3 to properly incorporate emissions originating from biomass burning, compared to that of CTM2. For this we compare our modelled CO in the troposphere with measurements carried out in 2005 by CARIBIC (Civil Aircraft for the Regular Investigation of the atmosphere Based on an Instrument Container, http://www.caribic-atmospheric.com, Brenninkmeijer et al., 2007). In the year 2005, most CARIBIC flights were operated between Europe and South America, and model results are interpolated on-line to the spatial and temporal locations of measurements. Figure 11 shows all CO measurements for all flights in 2005 (black), along with corresponding model output (CTM3 in red, CTM2 in green). Measurements located in the CTM stratosphere (defined in Sect. 3.3) are shown in blue. While observations reached up to ∼ 500 ppb, we have cut off the y-axis at 250 ppb to make the figure more readable; all model results are lower than 250 ppb. Both CTM2 and CTM3 produce CO remarkably close to measured values. Even though most of the flight measurements are carried out over the Atlantic, spikes originating from biomass burning events in August and October can be seen (Ebinghaus et al., 2007). These events are well modelled, which is due to the combination of realistic meteorological data and that the horizontal distribution of model CO emissions from biomass burning to a large degree is well represented. It should be noted that recently Hooghiemstra et al. (2012) found that GFEDv3 emissions are too low in South America for the years 2006–2010, which could also be the case for 2005 used in our study. CTM3 produce larger spikes than CTM2 does, due to the improved vertical transport in CTM3. It is expected that with a higher temporal resolution on GFEDv3 emissions, these spikes would be even higher.

3.3 Global diagnostics

Several global diagnostics are frequently used to evaluate atmospheric models, such as the CH$_4$ lifetime, the average tropospheric OH concentration, the O$_3$ burden and the mass flux of O$_3$ from the stratosphere into the troposphere. Here we present these and also the lifetime of N$_2$O.

The diagnostics are in general calculated within domains between the model surface and four different upper boundaries. These upper boundaries are our model tropopause (2.5 PVU), 200 hPa, the 150 ppb O$_3$ surface, and
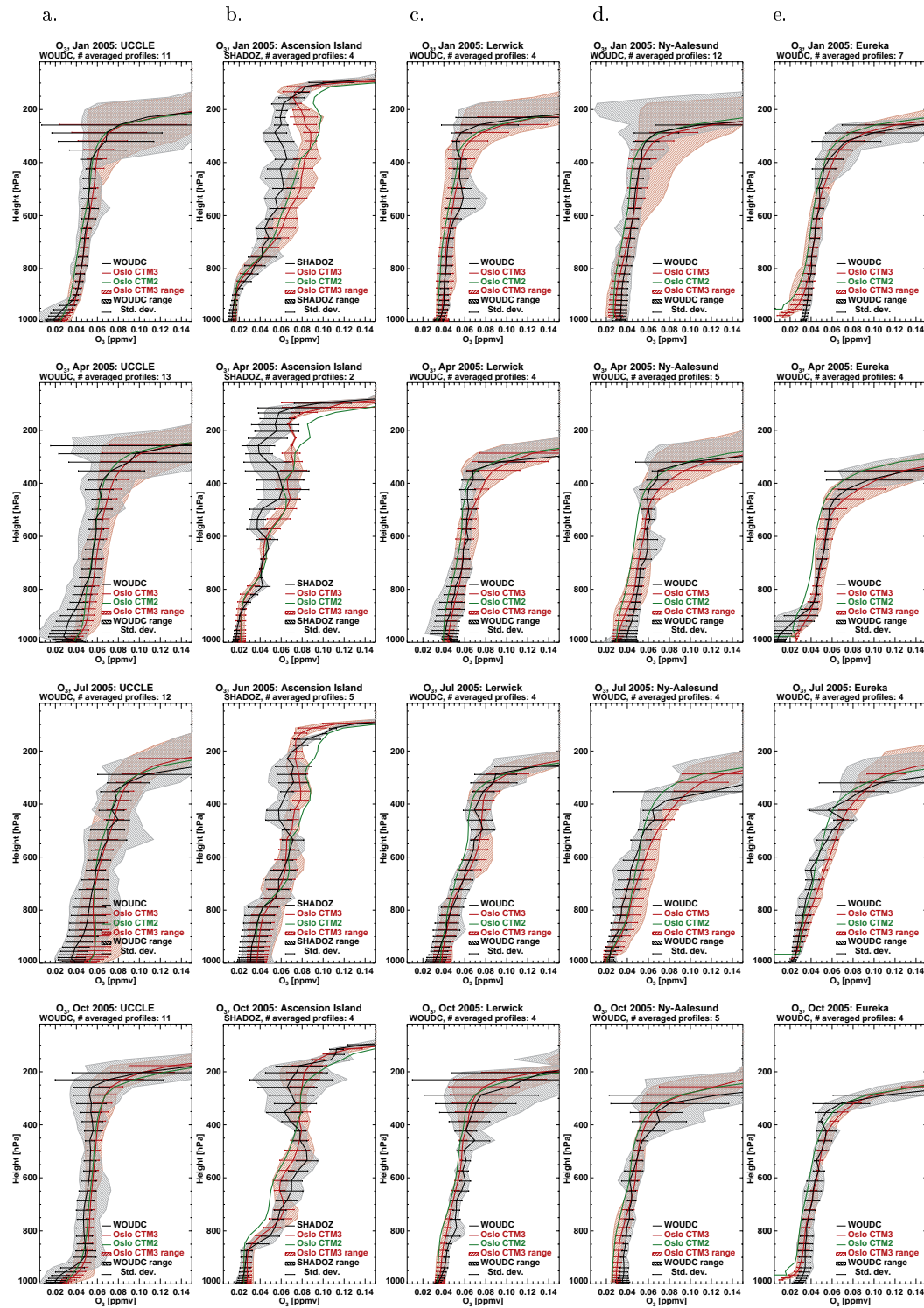
Fig. 10. Modelled (C3_ssn and C2_ssn) and measured monthly mean vertical profiles of O$_3$ at selected sonde stations **(a–e)** in 2005. Different months for each station are shown in columns.

the model top (LTOP). The prior defines the tropopause at potential vorticity of 2.5 PVU (Holton et al. (1995), 1 PVU is 10^{-6} K m^2 kg^{-1} s^{-1}), with an upper limit of 380 K poten-

tial temperature and a somewhat arbitrary lower limit of 5 km (only occurs occasionally). For 2005, the fraction of air below the CTM tropopause is 80.2 %, while below 200 hPa it
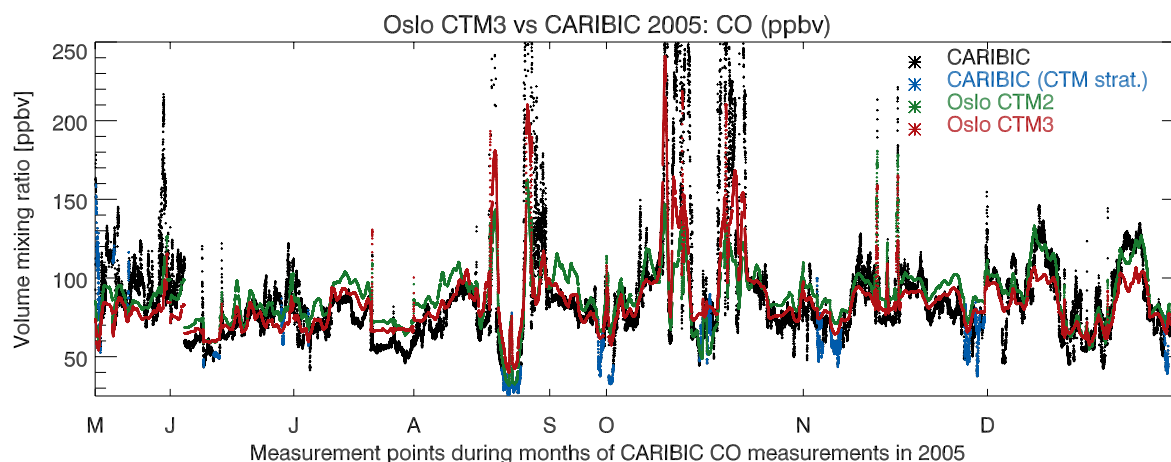
Fig. 11. Oslo CTM3 (red) and CTM2 (green) compared with all CO measurements from CARIBIC flights (black) in 2005, resulting in an asynchronous time axis where only the start of each month is shown. Measurements carried out at locations in the CTM stratosphere (defined by 2.5 PVU) are shown in the blue "CARIBIC (CTM strat.)"

is 79.7 %. The fraction below 150 ppb O_3, in general higher than 2.5 PVU, differs slightly for the different runs: 84.6 % for C3, 85.5 % for C3_ssn, 83.4 % for C2 and 83.5 % for C2_ssn.

3.3.1 OH concentration

OH is the main oxidizing species in the troposphere, driven by photolysis rates. It is usually referred to in context of CH_4 lifetime, which will be discussed in Sect. 3.3.2. Here we present average OH from the different year 2005 simulations, calculated from monthly means of OH, air mass, pressure (p) and temperature (T). Two averaging kernels are used to calculate OH average; first the OH concentration is weighted by air mass and by the loss rate to CH_4 ($\exp(-1775/T)$), and second it is weighted by air mass and a simplified loss rate of CO using pressure ($1 + 0.6p$). The values are presented in Table 3, showing small differences for the different model domains.

In simulation C3, the annual mean CH_4-loss weighted OH concentration is $\sim 1.55 \times 10^6$ molec cm^{-3}, while in C2 it is $\sim 1.35 \times 10^6$ molec cm^{-3}. C3_ssn and C2_ssn produce values of $\sim 1.39 \times 10^6$ molec cm^{-3} and $\sim 1.22 \times 10^6$ molec cm^{-3}, respectively. For the OH average weighted against ($1 + 0.6p$), the averages are about 10 % smaller, but otherwise show the similar picture. We find that the vertical distribution of lightning to a fixed height of 16 km, as in CTM2 (Sect. 2.6), can explain about 50 % of the larger OH in CTM3 compared to CTM2. Otherwise, the increase is partly due to updates in solar fluxes and O(^{1}D) quantum yields (Hsu and Prather, 2009) and to differences in wet scavenging. The latter will affect OH through scavenging of, e.g., HNO_3 and H_2O_2, and will be addressed in Sect. 3.4.

Inclusion of sulfur and nitrate modules in CTM2 and CTM3 is important for OH, reducing it by 10–11 %. As already noted, conversion of N_2O_5 to HNO_3 on aerosols is included in the tropospheric chemistry scheme; it is the same in CTM3 and CTM2, and cannot explain the reduction in OH. From a 3-yr simulation (1997–1999) where nitrate aerosols are excluded, we find that about 50 % of the reduced OH may be due to sulfur chemistry removing OH through oxidation processes. As important is the sequestering of NO_x in nitrate aerosols (Myhre et al., 2006), which with subsequent washout adds a removal process of atmospheric nitrogen and thus a source of OH. OH reduction due to the nitrate module is most pronounced in NH winter/spring, when the wind-driven production of sea salt is largest.

Lamarque et al. (2010) showed that the RETRO anthropogenic CO emissions are probably too low, and other studies indicate that estimated CO emissions probably are too low in general (Pison et al., 2009; Kopacz et al., 2010). However, when increasing the anthropogenic CO emissions by 20 % in CTM3, we find < 1 % reduction in the OH numbers. Locally, the accompanying effect on O_3 can be large (up to 100 %), but in the zonal mean O_3 increases by less than 1 %. Even though the RETRO emissions may be low on CO, OH is not very sensitive to an increase in anthropogenic emissions.

Further, fast-JX allows for letting aerosols affect photochemistry through scattering and absorption, thereby possibly changing OH. Bian et al. (2003) found the effect of aerosols on photolysis to reduce global OH by ~ 8 %, similar to Martin et al. (2003) who also found a larger reduction in the Northern Hemisphere than in the Southern Hemisphere. This will be investigated in the already-mentioned upcoming CTM3 paper focusing on aerosols.

Table 3. Annual average of OH in the atmosphere, calculated from monthly means of OH, air mass, temperature (T) and pressure (p, in bars), for the air mass below the model PVU-based tropopause height, air mass below 200 hPa, air mass below the modelled 150 ppb O_3 surface, and the air mass in the whole model domain (up to top level LTOP). The average OH concentration is weighted by air mass and by **(a)** the loss rate of CH_4 ($\exp(-1775/T)$) and **(b)** the approximate loss rate of CO ($1 + 0.6p$). All numbers are for the year 2005, and the ranges of monthly means are given in parentheses.

a. [OH] (10^6 molec cm^{-3}) weighted by air mass and $\exp(-1775/T)$.
Diagnosed between surface and selected upper boundaries.

Sim.	2.5 PVU	200 hPa	150 ppb O_3	LTOP
C3	1.55 (1.47–1.70)	1.54 (1.46–1.70)	1.54 (1.46–1.70)	1.55 (1.47–1.69)
C3_ssn	1.39 (1.30–1.56)	1.38 (1.29–1.56)	1.38 (1.29–1.56)	1.39 (1.31–1.56)
C2	1.35 (1.27–1.49)	1.34 (1.26–1.48)	1.35 (1.26–1.48)	1.35 (1.28–1.48)
C2_ssn	1.22 (1.11–1.36)	1.21 (1.11–1.36)	1.21 (1.11–1.36)	1.22 (1.13–1.37)

b. [OH] (10^6 molec cm^{-3}) weighted by air mass and ($1 + 0.6p$).
Diagnosed between surface and selected upper boundaries.

Sim.	2.5 PVU	200 hPa	150 ppb O_3	LTOP
C3	1.40 (1.33–1.54)	1.39 (1.32–1.53)	1.39 (1.32–1.53)	1.37 (1.30–1.50)
C3_ssn	1.26 (1.16–1.41)	1.24 (1.15–1.40)	1.25 (1.16–1.40)	1.23 (1.14–1.38)
C2	1.21 (1.13–1.34)	1.20 (1.12–1.32)	1.21 (1.13–1.33)	1.18 (1.11–1.31)
C2_ssn	1.09 (0.99–1.23)	1.07 (0.97–1.21)	1.08 (0.98–1.22)	1.06 (0.97–1.20)

Table 4. Annual average lifetimes of CH_4 (τ_{CH_4}), diagnosed as the total CH_4 burden (surface to model top) divided by the loss of CH_4 within the four diagnosed domains: below the model PVU-based tropopause height, below 200 hPa surface, below the modelled 150 ppb O_3 surface, and the whole model domain (up to top level LTOP). All numbers are for the year 2005, and the ranges of monthly means are given in parentheses.

τ_{CH_4} (yr). Diagnosed between surface and selected upper boundaries.

Sim.	2.5 PVU	200 hPa	150 ppb O_3	LTOP
C3	7.98 (6.96–8.62)	8.03 (7.02–8.67)	7.94 (6.94–8.56)	7.55 (6.64–8.11)
C3_ssn	8.86 (7.59–8.83)	8.89 (7.63–9.85)	8.81 (7.56–9.75)	8.33 (7.22–9.16)
C2	9.18 (7.98–10.06)	9.26 (8.06–10.15)	9.14 (7.95–10.01)	8.64 (7.59–9.39)
C2_ssn	10.20 (8.68–11.43)	10.31 (8.77–11.55)	10.16 (8.65–11.37)	9.54 (8.22–10.58)

3.3.2 CH_4 and N_2O lifetimes

The main atmospheric sink of the greenhouse gas CH_4 is OH, predominantly in the troposphere. CH_4 is also lost to Cl and $O(^1D)$, mainly in the stratosphere, with small additional loss to Cl in the marine boundary layer. In the troposphere, CTM2 and CTM3 neglect Cl chemistry and also the CH_4 loss to $O(^1D)$. Instead of simulating surface emissions and soil uptake, both models use prescribed surface concentrations of CH_4, based on CTM2 simulations with CH_4 emissions and surface deposition (HYMN-report, 2010).

Table 4 lists global mean CH_4 lifetimes in 2005, calculated from monthly mean total CH_4 burden and cumulative monthly CH_4 loss. On a monthly basis, CH_4 lifetime is shortest during NH summer, consistent with the seasonal cycle of OH. Our three tropospheric definitions produce lifetimes that differ by up to 10 %, with the shortest lifetimes estimated using the 150 ppb O_3 boundary because it includes the tropical upper troposphere and some high latitude stratospheric loss

due to Cl and $O(^1D)$. We use the O_3-based tropopause definition in this section.

The tropospheric CH_4 lifetime in CTM3 is 7.9 yr, versus 9.1 yr in CTM2, consistent with the larger OH concentration in the updated model. Adding sulfur and nitrate aerosols increases the lifetime by ~ 12 %, to 8.9 yr in CTM3 (sim. C3_ssn). This aerosol effect is slightly smaller in CTM2 (~ 11 %). Past studies have found a similar impact of aerosols on CH_4 lifetime via surface reactions (Martin et al., 2003; Macintyre and Evans, 2010). Recent evaluation of the CH_4 lifetime (Prather et al., 2012), total and against loss by tropospheric OH, take into account the recently observed CH_3CCl_3 decay (Montzka et al., 2011) and up-to-date estimates with uncertainties of the other factors that go into the lifetime calculation. Prather et al. (2012) report the CH_4 lifetime with respect to tropospheric OH (the so-called tropospheric OH-lifetime) to be 11.2 ± 1.3 yr, which is longer than earlier estimates of (8.7 ± 1.3 yr IPCC AR4, 2007). Thus, the tropospheric OH-lifetime of CH_4 in CTM3, when including

Table 5. Modelled annual average O_3 burden [Tg (O_3)] for different definitions of tropospheric domains.

| | Diagnosed between surface and a selection of upper boundaries. | | | | | |
| | 2.5 PVU | | 200 hPa | | 150 ppb O_3 | |
Simulation	2005	1998–2005	2005	1998–2005	2005	1998–2005
C3	360 (340–386)	360	356 (342–374)	357	391 (373–415)	390
C3_ssn	331 (310–359)	331	344 (327–361)	345	363 (344–388)	362
C2	352 (336–376)	354	344 (331–355)	347	372 (358–389)	373
C2_ssn	326 (309–352)	328	317 (303–331)	321	346 (332–366)	347

aerosols, is shorter than the recent estimate, but within the range of older estimates. Adding aerosols to the photolysis calculation, currently neglected in CTM2 and CTM3, could increase the tropospheric-OH lifetime by $\sim 8\%$ (Bian et al., 2003; Martin et al., 2003) to ~ 10 yr in CTM3 and bring CTM3 within the recent methyl chloroform constraint.

Over the entire atmosphere, the CH_4 lifetime due to all chemical losses is 9.5 yr in CTM2 and 8.4 yr in CTM3 in the simulations with aerosols. If we include estimates of the soil and tropospheric Cl sinks (Prather et al., 2012), the total lifetimes decrease 7–10 %. Because the CTMs use prescribed surface CH_4 concentrations, but have different CH_4 lifetimes, they calculate different CH_4 burdens. These are 4830 Tg and 4822 Tg for simulations C2 and C3, respectively (4831 Tg and 4825 Tg for C2_ssn and C3_ssn, respectively). All are about 2 % lower than the burden estimated by IPCC (4932 Tg, IPCC AR4, 2007). The stratospheric lifetime of CH_4 is 147 yr and 141 yr in CTM2 and CTM3, respectively. These lifetimes change only slightly ($< 3\%$) after including tropospheric aerosols. We estimate that OH, Cl, and O(^{1}D) cause 47 %, 29 % and 24 % of stratospheric CH_4 oxidation, respectively.

Another important greenhouse gas is N_2O, which in general is inert in the troposphere, while having a small stratospheric loss. Hence, N_2O has a long lifetime, and in CTM3 we calculate it to be 144 yr, with a monthly average range of 133–155 yr. Because the photolysis code used in CTM2 (fast-J2) had smaller solar fluxes in the 200–230 nm region where N_2O and O_2 are photolyzed in most of the stratosphere, the N_2O lifetime in CTM2 is longer, 151 yr. It is worth noting that the CTM2 and CTM3 keep N_2O fixed at the surface and at the uppermost model layer – with the same mixing ratios in both models. The N_2O lifetimes are probably too long because of slow vertical transport in the cycle 36r1 meteorology used in both calculations, somewhat evident in the underestimate of the CTM N_2O profiles (Fig. 7).

3.3.3 O_3 burden

We have calculated the tropospheric burden of O_3 using the tropospheric domains defined in Sect. 3.3.2. The numbers are presented in Table 5 for the year 2005, and as the mean of years 1998 to 2005. Also, the ranges of the monthly mean values for 2005 are listed. The interannual variation is small,

as are the differences between CTM3 and CTM2. Below the 150 ppb O_3 surface, modelled burden in 2005 is 391 Tg for C3 and 363 Tg for C3_ssn, while the Oslo CTM2 produces 372 Tg for C2 and 346 Tg for C2_ssn. Similar values are found for the 1998–2005 averages. Within the domains below our model tropopause and below 200 hPa, which are smaller domains, the O_3 burden is 5–10 % lower than when using the 150 ppb O_3 definition. Based on several model studies, Stevenson et al. (2006) presented a tropospheric burden range of 344 ± 39 Tg, calculated from all air masses with O_3 less than 150 ppb. The C3_ssn and C2_ssn runs lie well within this range, with values of 346 and 363 Tg, respectively. When excluding the aerosols our models lie in or just above the upper end of the range. It should be noted that Stevenson et al. (2006) used different emissions, which could somewhat affect the comparison.

By distributing L-NO_x vertically to a fixed altitude of 16 km in the CTM3, we find that the higher burden in CTM3 compared to CTM2 is most likely due to the differences in L-NO_x vertical scaling. Otherwise, as noted in Sect. 3.3.2, the main differences between CTM2 and CTM3 are due to the new fast-JX photochemistry being somewhat more active, and differences in wet scavenging. While this produces higher O_3 burden in CTM3, it does not change the relative effect of nitrate aerosols when compared to CTM2, about 8 %. A halving of the operator-split time step (C3_1/2) only changes the burden by $\sim 2\%$, indicating that the tropospheric O_3 burden is not very sensitive to the operator-split time step.

3.3.4 Cross-tropopause O_3 flux

The stratosphere to troposphere exchange (STE) is the net flux from the stratosphere into the troposphere, and is estimated based on observations to be 540 ± 140 Tg (O_3) yr^{-1} (Murphy and Fahey, 1994; Gettelman et al., 1997; IPCC AR4, 2007). Model studies suggest STE around 500 Tg (O_3) yr^{-1} (McLinden et al., 2000; Olsen et al., 2001; Hsu et al., 2005), and Stevenson et al. (2006) reported a model range of 552 ± 168 Tg (O_3) yr^{-1}.

In Oslo CTM3 the STE calculation follows Hsu et al. (2005), which is the same method as used by Stevenson et al. (2006) (referred to as S2006 below). Due to technical challenges, this has not been possible in CTM2, and while

Søvde et al. (2011a) presented CTM2 STE, their method is not comparable to the CTM3 treatment – hence we present only CTM3 in this study. During a certain time period in CTM3 (e.g., a month), STE of O_3 is calculated as a residual of column mass budgets between the model surface and a certain O_3 isopleth (Hsu et al., 2005). Below the given O_3 isopleth and for the given time period, this can be written

$$\left(\frac{dM}{dt}\right)_{\text{tot}} = \left(\frac{dM}{dt}\right)_{\text{chem}} - S + F_{s\to t} - F_{h,t\to t}, \qquad (7)$$

where $\left(\frac{dM}{dt}\right)_{\text{tot}}$ is the total change in O_3, $\left(\frac{dM}{dt}\right)_{\text{chem}}$ is the corresponding chemical tendency of O_3 (corresponding to P−L in S2006), S is sink processes such as dry deposition at the surface and wet scavenging (D in S2006), $F_{h,t\to t}$ is the horizontal flux of O_3 out of the column (troposphere to troposphere), and $F_{s\to t}$ is the STE to be inferred (both horizontal and vertical). Hence, the left-hand side of Eq. (7) is diagnosed as the change from the beginning to the end of the diagnosed time period, while the right-hand side terms are summed up for all time steps during the time period. S2006 assumed no tropospheric trend, i.e. $\left(\frac{dM}{dt}\right)_{\text{tot}} + F_{h,t\to t} = 0$ in Eq. (7).

We have found the 150 ppb isopleth of O_3 used in S2006 to confuse stratospheric air with tropospheric air for O_3-hole conditions, hence we have lowered this isopleth to 120 ppb to minimise this problem. For the year 2005, Oslo CTM3 produce an annual STE of 296 Tg (O_3) yr^{-1} for C3 and 300 Tg (O_3) yr^{-1} for C3_ssn. The aerosol included in the ssn-runs are not calculated in the stratosphere, but changes the 120 ppb isopleth slightly, thereby changing STE slightly. On a monthly basis, the STE is largest in NH summer, while smallest in NH winter. In an atmosphere without polar stratospheric clouds (PSCs), the 120 ppb isopleth can be very different than when PSCs are included, and stratospheric burden of O_3 is higher. Hsu and Prather (2009) found PSCs to reduce STE by 97 Tg (O_3) yr^{-1}, and when testing this in CTM3 we find the STE to increase by a comparable amount to 427 Tg (O_3) yr^{-1} for the year 2005. It may be that STE inferred from observations does not fully take the PSC-O_3 hole into account, so modelled STE without PSCs provide insight into this. However, it should be kept in mind that the 120 ppb isopleth has changed.

With the C3_1/2 simulation, STE is slightly reduced to 288 Tg (O_3) yr^{-1}, due to a slightly changing O_3 isopleth and less overshooting of the stratospheric jet. We have also carried out a 14 month simulation starting on 1 November 2004, using an older cycle of the ECMWF IFS meteorology (cycle 29). With this, the Oslo CTM3 produce a STE of 353 Tg (O_3) yr^{-1}. While this higher flux indicates that the old cycle may provide a more realistic circulation, it should be noted that the N_2O FrIAC presented in Sect. 3.1.3 is much less pronounced for cycle 29 meteorological data, indicating that cycle 29 polar cap winds are less representative. Still, our comparison with MLS indicates that cycle 36r1 may have

too slow Brewer-Dobson circulation, which will give too low STE. Also, UCI-CTM calculations of STE (Hsu and Prather, 2009) are considerably reduced when using cycle 36r1.

Due to the difficulties with a calculated O_3 isopleth, such as stratospheric air being counted as tropospheric air for low O_3 and tropospheric O_3-rich air in polluted areas being counted as stratospheric air, we have also tested calculating STE through an alternative surface, namely the troposphere defined by the artificial e90-tracer (Prather et al., 2011). With this we find the annual STE of 285 Tg (O_3) yr^{-1} for C3_ssn close to the value from using the 120 ppb isopleth. Also, the reduction when halving the operator-split time step is similar to what was found using the 120 ppb isopleth.

3.4 Wet scavenging

Wet removal of atmospheric HNO_3 and H_2O_2 is important for the NO_x and OH abundances. Less removal of H_2O_2 will increase tropospheric OH, and less removal of HNO_3 will increase atmospheric NO_x. HNO_3 can be recycled back to NO_x by either photodissociation or reaction with OH, but most of it is lost to the surface by wet or dry scavenging. In addition, NO_x will by itself change OH by shifting the HO_2/OH balance.

Hence, if CTM3 wet scavenging is less efficient than for CTM2, the amount of OH and NO_x will be higher than in CTM2. Due to the complexity of the coupled system, it has not been possible to separate out the contributions to OH from increased NO_x due to the wet scavenging scheme from that due to the update to fast-JX. It should be noted that a major difference from CTM2 to CTM3 scavenging is that CTM2 treats all large-scale precipitation as liquid rain, while CTM3 separates it into ice and liquid precipitation. The only species removed by ice in CTM3 is HNO_3, and even that is reduced relative to liquid rain (CTM2). The convective scavenging scheme is the same in both, and treats all convective precipitation as convective rain.

To compare the wet scavenging in CTM3 to CTM2, we have carried out idealised 7-day simulations with only transport and wet removal of HNO_3 and H_2O_2. Tracer removals by large-scale and convective precipitation are diagnosed hourly. We present results for 1–7 January 2005, but find similar results for other months. The idealised tracers are initialised from the C3_ssn simulation at 1 January 2005.

In the upper troposphere, CTM3 scavenges less by rain than CTM2, while in the lower tropical troposphere, more is scavenged in CTM3. In total HNO_3 burden, CTM3 removes less HNO_3 than does CTM2; starting from a tropospheric burden of 1.43 Tg (HNO_3), CTM3 ends at 0.62 Tg (HNO_3) after 7 days, while CTM2 ends at 0.40 Tg (HNO_3). The higher HNO_3 in CTM3 is due to large-scale precipitation being constrained by cloud fraction and hence less effective; the amount removed at each time step is about 30 to 40 % lower than for CTM2. Large changes in the initial scavenging rate occur because the initialized fields (C3_ssn) were

inconsistent with the rapid large-scale scavenging in the upper troposphere by CTM2. By looking at the decay rates over the last few days of the experiment, when the HNO_3 distribution more closely matches each CTM's pattern of scavenging, we find that for CTM3 the effective tropospheric lifetime of HNO_3 is 9.3 days and for CTM2 it is 6.9 days, confirming the more rapid loss in CTM2. Spatially we find that CTM3 removes less in the upper troposphere, while removing slightly more in the lowermost tropical troposphere, similar to results in Neu and Prather (2012).

The amount removed by convective scavenging is initially similar in CTM2 and CTM3, however, after a few time steps CTM3 removes more by convective scavenging than CTM2, compensating for the reduced large-scale scavenging. However, the fraction of tropospheric HNO_3 washed out by convective rain at each time step remains similar in CTM2 and CTM3, since the process is unchanged in the CTM update.

Less effective wet scavenging is also found for H_2O_2 in CTM3. The tropospheric H_2O_2 burden starts at 3.86 Tg, and after 7 days of only transport and wet scavenging, CTM3 ends at 1.66 Tg, while CTM2 ends at 0.95 Tg. After 7 days these correspond to a tropospheric lifetime of 5.9 days for CTM2 and 11.4 days for CTM3. Initially, CTM2 removes about 4 times as much as CTM3 in large-scale scavenging, which is a larger difference than for HNO_3, because H_2O_2 is not removed by ice in CTM3 but is removed in CTM2. The fraction of the tropospheric burden removed by convective scavenging is smaller in CTM3, which was not the case for HNO_3. While Neu and Prather (2012) found H_2O_2 uptake in ice to have a very small effect on O_3, the differences from CTM2 to CTM3 may be large enough to re-examine this result.

Our chemistry simulations C2 and C3 show similar results; less HNO_3 is scavenged in CTM3, giving the latter a higher tropospheric burden of HNO_3. The very different pattern of scavenging from CTM2 to CTM3 also contributes to this by changing the relative distribution of HNO_3 throughout the troposphere. In the ssn-simulations, we see a reduction of tropospheric HNO_3 by about 28 % in CTM2 and 32 % in CTM3. This agrees well with what we found for OH and lifetime of CH_4. The slightly higher effect on HNO_3 in CTM3 is probably due to the less efficient wet scavenging in CTM3 also affecting sea salt and nitrate aerosols, allowing more HNO_3 to be bound in nitrate particles. Also important are differences in other loss processes such as photodissociation and the loss to OH.

The higher HNO_3 burden in CTM3 is not distributed evenly. In the tropics, below 400 hPa, HNO_3 is about 50 % lower in CTM3, while in the extra-tropics CTM3 has up to twice as much HNO_3 as in CTM2, consistent with less scavenging. Inclusion of tropospheric aerosols reduces extra-tropical HNO_3 by up to 100 % in both CTM2 and CTM3, slightly more pronounced in CTM3. Because of the less efficient scavenging of sea salt aerosols in CTM3, more HNO_3 is taken up to form nitrate aerosols, explaining this difference.

Above 200 hPa, in the stratosphere, CTM3 has about half of the HNO_3 in CTM2, except in the upper stratosphere, where CTM3 again has somewhat higher HNO_3.

The higher OH in CTM3 than in CTM2 is partly due to less removal of H_2O_2; the CTM3 burden of H_2O_2 is ~ 12 % higher than for CTM2. Spatially, this difference comes from a decrease of ~ 20 % in the tropical troposphere, but an increase of up to 400 % in the upper troposphere due to lack of scavenging by ice. The more active photochemistry in fast-JX may dissociate more HNO_3 to NO_x, which again contributes to increasing OH and thereby to HNO_3 loss, but as already mentioned, it has not been possible to separate the effects of differences in scavenging and photochemistry.

3.5 Transport sensitivity studies

In order to document the differences in transport schemes between CTM2 and CTM3, we have carried out several 14-month sensitivity studies, listed in Table 2. Here we summarize the findings of shortening the operator-split time step and of using the more accurate polar cap transport in C3_pole.

When halving the CTM3 operator-split time step to 30 min (C3_1/2), the computing time is increased by almost 100 % because all processes are done twice. In general there are small differences between C3 and C3_1/2, except that the SH polar vortex edge seems to be better resolved in C3_1/2, with a sharper gradient across the vortex edge. Total O_3-hole column values at the polar side of the SH vortex edge are reduced by up to 20 %, while on the mid-latitude side the column values are increased by up to 5 %. These relatively large changes are only found across the sharp gradients, corresponding well with what was seen for age of air. A 15 min operator-split time step, compared to 60 min, increases these numbers to 25 % and 9 %, respectively, showing rapid convergence to CTM2, as was shown in Sect. 3.1.2. A similar effect can also be seen in the Northern Hemisphere, however to a smaller extent. The largest differences occur at mid to high latitudes between 300 hPa and 20 hPa, in O_3-hole conditions with strong asymmetric vorticies. Here, as we have discussed, the jets may overshoot. When halving the operator-split time step, the maximum increase outside the vortex and the maximum decrease inside the vortex are of the same order: In SH, this maximum change from C3 to C3_1/2 is ~ 25 %, enhanced further by 10–15 % for C3_1/4. In NH these numbers are smaller ($< 1/2$).

A halving of only the horizontal transport time step has also been tested, and found to produce somewhat sharper vortex gradients, although not to the same extent as for C3_1/2. This indicates that vertical transport is important for a better resolved vortex edge.

Using the more accurate polar transport (C3_pole) increases the computing time by about 50 %. Hence, not all model time steps require shorter transport time steps than for C3. The changes from C3 are small, and generally confined to the polar cap. However, the shorter transport time steps

produce a somewhat sharper gradient across the vortex edge than for C3. C3_pole may impose a shorter global transport time step, but keeps the operator-split time step unchanged. Thus, transport will essentially be similar to C3_1/2 when the transport time step is reduced, but still carried out for 60 min before chemistry is calculated. However, if the global transport time step does not need to change, it may be that the internal transport step (1-D pipe flow) requires a shorter time step, but this will not produce large differences between C3 and C3_pole. Again, this was also seen for the age of air studies.

Studies involving stratospheric polar cap transport in T42 horizontal resolution ($\sim 2.8°$), e.g., Arctic and Antarctic O_3 loss studies, will clearly benefit from using the C3_pole version combined with a shorter operator-split time step (e.g., C3_1/2), depending on the required accuracy at the vortex. When reduced accuracy at the vortex gradients is acceptable, e.g., when computational costs would otherwise prevent a calculation, the standard setup (C3) should be used. However, at higher horizontal resolutions, e. g. T159, the need of C3_pole may be reduced because C3 will have better effective polar resolution ($\sim 2.2°$) and accuracy than C3_pole at T42.

3.6 Computational efficiency

Part of the motivation for updating the Oslo CTM2 was to make the transport faster, allowing long or complex chemistry simulations. To determine how much faster the Oslo CTM3 is than CTM2, and how well they scale to the number of CPUs used, we have carried out tests for the T42L60 dataset used. Both models are run for one month (January 2005), a month which was found to represent model transport times well, on an Intel Xeon with 32 X7560 2.27 GHz CPUs.

While the CTM2 transport is parallelized over number of tracers, CTM3 is parallelized over pre-defined horizontal blocks, except for the horizontal transport which is parallelized over model vertical layers. This allows the CTM3 to be run in parallel even for one tracer, and because most processes can be treated in horizontal blocks, matching the number of blocks to a multiple of the number of CPUs allows for a more efficient parallel code. This is an important improvement, making the CTM3 more flexible for using a higher number of CPUs.

Full chemistry simulations with CTM2 and CTM3 (C3) were timed for T42L60 resolution, and CTM3 was found to be about 40 % faster than CTM2. Pure transport tests reveal closer to 50 % improvement, indicating that processes related to chemistry and diagnostics, which are now more realistic, take more time in Oslo CTM3 than in CTM2. The main contributors to this are the stratospheric microphysics, which is treated with a time step of 15 min instead of one hour in CTM2, but to some extent also the fast-JX cloud overlap treatment, as described in Sect. 2.4.

Both models scale rather well with number of CPUs; a doubling of CPUs generally makes the runs ~ 40 % faster. However, the gain is reduced dramatically for the CTM2 full chemistry run when shifting from 16 to 32 CPUs due to the number of transported tracers badly matching the numbers of CPUs. CTM3 shows only a slightly lower improvement from 16 to 32 CPUs than from 8 to 16 CPUs. Increasing the number of CPUs also makes the fractional time spent on non-parallelized processes larger, which could also explain some of the inefficiency of increasing CPU numbers.

Pure transport simulations are also carried out in a higher horizontal resolution, namely T159L60 ($1.125° \times 1.125°$ degrees), showing a larger reduction (~ 60 %) in computing time from CTM2 to CTM3 than at T42L60 resolution. Hence, high resolution simulations have become more feasible with the Oslo CTM3.

4 Discussion and conclusions

We have presented the chemical transport model Oslo CTM3 and documented the changes from the earlier version Oslo CTM2. The Oslo CTM3 provides a large improvement of the Oslo CTM2, due to updated and more physically based parameterizations, cleaner code, and increased computational efficiency due to the new structure. The latter makes the Oslo CTM3 better suited for parallelization and up to 40 % faster, enabling much higher spatial resolution or multi-year ensemble/sensitivity simulations. CTM3 also has improved the representation of sub-grid processes such as convection, scavenging, and fractional cloud cover. Thorough evaluation of the Oslo CTM3 against the Oslo CTM2 and against measurements shows that the new version is as good as, if not better than, the old model.

Oslo CTM3 allows for stable, large time steps for advection, and this is essential in enabling more complex (e.g., fully coupled stratosphere-troposphere, gas-aerosol) CTM simulations. Nevertheless, the larger time steps result in larger errors, particularly near the jets where a curving flow may partly overshoot (Prather et al., 2008). When necessary, these errors can be reduced by using a shorter operator-split time step (or advection time step) in the Oslo CTM3.

Stratospheric age of air compare well between CTM2 and CTM3. The oldest age of air in CTM3 is about 5.8 yr, slightly lower than in CTM2 (5.9 yr). The spatial distributions are similar in CTM2 and CTM3, except close to the Southern Hemisphere polar vortex edge. We have shown that reducing the transport time step causes CTM3 to converge to CTM2 results at equivalent time steps.

In general, Oslo CTM3 represents the stratosphere better than CTM2 due to the update to fast-JX. Total O_3 columns and burdens compare well with measurements and other model studies. In the troposphere there are some model differences due to differences in photochemical reaction rates and also due to different schemes for wet scavenging by rain.

The CTM3 wet scavenging scheme is more physically based than the old version, although it removes 20–50 % less of soluble species from the troposphere. CTM3 separates large-scale scavenging by ice and rain, but the only species subject to ice scavenging is HNO_3. Because CTM2 assumes all precipitation to be rain, it generally scavenges more than CTM3 does. As a consequence, with only the tropospheric and stratospheric chemistry schemes included, NO_x is somewhat increased and OH is slightly high, giving a corresponding shorter CH_4 lifetime with respect to OH (8 yr). Including sulfur chemistry and nitrate aerosols increases the CH_4-lifetime by ~ 12 % to 8.9 yr. We find that increasing anthropogenic CO emissions in the RETRO dataset by 20 % cannot reduce OH in CTM3 by more than ~ 1 %, and does not affect the CH_4-lifetime much. With the improvement to fast-JX and the new scavenging in CTM3, it may be that CTM2 has a longer CH_4-lifetime for the wrong reasons, e.g., by distributing lightning NOx incorrectly.

The new stratosphere to troposphere flux calculation through the 120 ppb O_3 surface yields a somewhat low value of $\sim 300\,\mathrm{Tg}\,(O_3)\,\mathrm{yr}^{-1}$ for the year 2005. Halving the transport time step reduces the STE slightly. Using an older cycle of ECMWF meteorological data yields a higher flux more consistent with the semi-empirical STE fluxes of Murphy and Fahey (1994) and Olsen et al. (2001). This, and the model underestimate of N_2O compared with satellite observations, indicate that the ECMWF IFS cycle 36r1 has a degraded performance in the stratosphere with a too slow Brewer-Dobson circulation, especially for tropical upwelling. It has not been possible to test the newer cycle 38. We have also shown that turning off PSC heterogeneous chemistry increases the calculated STE, partly due to changing of the 120 ppb surface. A new surface for calculating STE has been introduced, namely the tropopause defined by the e90-tracer (Prather et al., 2011), which produces STE close to the STE using the 120 ppb O_3 isopleth.

Two polar cap transport treatments are available, and when studying polar cap transport care should be taken to use the most accurate method. Reducing the transport time step results in sharper polar vortex edges, and should be considered when studying processes close to or inside polar vortices.

It should be mentioned that the Oslo CTM3 also can be run without stratospheric chemistry, following a similar approach as in Skeie et al. (2011), using a model climatology for stratospheric O_3 and NO_x from a simulation with tropospheric and stratospheric chemistry. This would make the simulations faster due to fewer transported species, and may be an option for studies of the lower to middle troposphere, where the importance of the stratosphere is small.

Supplementary material related to this article is available online at: http://www.geosci-model-dev.net/5/ 1441/2012/gmd-5-1441-2012-supplement.pdf.

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